

Holographic flavor in $\mathcal{N} = 4$ gauge theories in 3d from wrapped branes

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Holographic flavor in $\mathcal{N} = 4$ gauge theories in 3d from wrapped branes

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ABSTRACT: We study the addition of flavor to the gravity dual of $\mathcal{N} = 4$ three-dimensional gauge theories obtained by wrapping N_c D4-branes on a two-cycle of a non-compact Calabi-Yau two-fold. In this setup the flavor is introduced by adding another set of D4-branes that are extended along the non-compact directions of the Calabi-Yau which are normal to the cycle which the color branes wrap. The analysis is performed both in the quenched and unquenched approximations. In this latter case we compute the backreacted metric and we show that it reproduces the running of the gauge coupling. The meson spectrum and the behavior of Wilson loops are also discussed and the holographic realization of the Higgs branch is analyzed. Other aspects of this system studied are the entanglement entropy and the non-relativistic version of our backgrounds.

KEYWORDS: Brane Dynamics in Gauge Theories, Gauge-gravity correspondence, D-branes, AdS-CFT Correspondence.

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1. Introduction

The gauge/gravity correspondence [1] is one of the major achievements of string theory in the last ten years (see [2] for a review). Its extension to include more realistic theories is, clearly, a topic of great interest. In particular, the addition of matter degrees of freedom is essential to get nearer to the string theory description of QCD-like theories. This addition can be performed by including extra (flavor) branes [3], which should be extended along all the gauge theory directions and wrap a non-compact cycle in order to make its worldvolume symmetry a global symmetry from the point of view of the gauge theory living in the color branes. They introduce an open string sector that corresponds to having hypermultiplets transforming in the fundamental representation in the gauge theory side.

If the number of flavors, N_f , is small compared with the number of colors, N_c , one can treat the flavor branes as probes in the supergravity background created by the color branes. This defines the so-called quenched approximation, which corresponds, in the field theory side, to neglecting quark loops, which are suppressed by factors $1/N_c$ in the 't Hooft large N_c expansion [4]. This quenched holographic flavor has been explored extensively in the past few years. In particular, by analyzing the normalizable fluctuations of the probe branes, the spectra of mesonic excitations of different theories have been analyzed (see [5] for a review and a list of references).

On the contrary, if the number of flavors is of the order of the number of colors ($N_f \sim N_c$) the backreaction of the flavor branes on the geometry cannot be ignored and one has to deal with a system of gravity plus branes, the latter acting as dynamical sources for the different supergravity fields. On the field theory side the inclusion of the backreaction in this $N_f \sim N_c$ regime is equivalent to considering the so-called Veneziano limit [6], in which N_c and N_f are large and their ratio N_f/N_c is fixed. In this limit quark loops are not suppressed and the flavor is unquenched.

Another direction in which the original gauge/gravity duality has been generalized is by extending it to theories with lower amounts of supersymmetry. A general strategy to carry out this extension of the correspondence to less supersymmetric models is to consider higher dimensional branes wrapping cycles. At energies small compared with the size of the cycle the theory becomes effectively four-dimensional. Moreover, the gauge theory living on the worldvolume of the wrapped brane has to be topologically twisted in order to preserve some fraction of supersymmetry. Examples of duals of $4d$ gauge theories constructed in this way are the geometries analyzed in [7] for $\mathcal{N} = 1$ and in [8] for $\mathcal{N} = 2$, which correspond to D5-branes wrapping a two-cycle inside a Calabi-Yau (CY) manifold.

In this paper we will study the dual of $\mathcal{N} = 4$ gauge theories in three space-time dimensions, obtained by wrapping D4-branes of the type IIA theory on a two-cycle of a Calabi-Yau two-fold. The corresponding unflavored supergravity solution was found in [9] and studied in detail in [10], where it was shown to reproduce the exact perturbative running coupling constant and the metric of the moduli space of the gauge theory. In the present paper we will analyze the addition of flavor branes in this setup. These flavor branes are also D4-branes, which are extended along non-compact directions of the CY two-fold in such a way that no further supersymmetry is broken.

We will analyze the addition of flavor to the $3d$, $\mathcal{N} = 4$ theory both in the quenched and unquenched approaches. In the quenched formalism we will compute numerically the spectrum of mesonic excitations and we will be able to find a very good analytical estimate of the masses. We will also study the Wilson loops and the corresponding quark-antiquark potentials. In this latter case we will discover that a phase transition is produced when some of the parameters of the solution are varied and we will compute the corresponding critical exponents.

We will include the backreaction by following the proposal of ref. [11], in which the localized brane sources are substituted by a continuous distribution of flavor branes in their transverse directions (see also [12] for a similar analysis in the context of non-critical string theory). This smearing brane approach has been applied successfully to various string duals in several dimensions ([13]–[25]). It has the advantage that it avoids having the δ -function sources of the localized branes in the BPS equations and, therefore, makes searching for solutions more feasible.

In our case we will start with the same ansatz for the metric as in the unflavored system. However, the presence of flavor branes modifies the Bianchi identity of the RR field strength F_4 , which forces us to modify the unflavored ansatz for F_4 . Once this fact is taken into account, one can get a system of first-order differential equations by imposing that the supersymmetric variation of the gravitino and dilatino of type IIA supergravity vanish. These first-order equations are simple, but they are difficult to solve in general. However, we will see that they can be solved analytically in certain regions of the space. Remarkably, this is enough to compute the modification of the running of the coupling constant due to the matter hypermultiplets in the gravity solution. We will show that this modification matches exactly the field theory results.

The organization of the rest of this paper is the following. We will start in section 2 by reviewing the unflavored background. We will set up our notations and write the BPS equations and their solution for the unflavored system. In section 3 we study the addition of flavor branes. In particular, we obtain the BPS equations for the backreacted background, which are then numerically integrated. Section 4 is devoted to analyzing the matching between our flavored background and its field theory dual. By means of a probe calculation we verify that the running of the gauge coupling with flavor is reproduced by our solution. We also discuss in this section how to realize the Higgs branch in our setup.

The analysis of the meson spectrum is the object of section 5. We study the mass levels both in the quenched and unquenched solutions and we also consider the meson spectra in the Higgs branch. In section 6 we explore the behavior of the Wilson loops. In our study of the energy of the quark-antiquark pair we will find some critical phenomena and we will evaluate the corresponding critical exponents. Section 7 is devoted to presenting our conclusions and summarizing our results.

The paper is completed with several appendices. In appendix A we give details of the derivation of the BPS equations for the general flavored system and we check that the equations of motion of the system are satisfied if the BPS equations hold. In appendix B we find the supersymmetric embeddings of D4-brane probes in our background. In appendix C we find additional solutions of the unflavored equations and we obtain the background dual

to a non-relativistic system that can be generated by deforming our solutions. Finally, in appendix C we analyze the entanglement entropy for our model in the UV.

2. The unflavored background

Following the analysis of refs. [9] and [10], let us consider the background of type IIA supergravity created by a stack of N_c D4-branes wrapped on a two-cycle \mathcal{C}_2 of a Calabi-Yau cone of complex dimension two, according to the following brane setup

	$\overbrace{\hspace{10em}}^{\text{CY}_2}$									
	$\mathbb{R}^{1,2}$			S^2		N_2		\mathbb{R}^3		
D4	-	-	-	○	○	·	·	·	·	·

where S^2 represents the directions of the two-cycle (which is topologically a two-sphere) and N_2 are the directions of the normal bundle to \mathcal{C}_2 . In the above setup a circle represents wrapped directions, whereas the symbols “-” and “.” denote unwrapped worldvolume and transverse directions respectively. We shall parameterize the cycle \mathcal{C}_2 by means of two angular coordinates $(\tilde{\theta}, \tilde{\phi})$ with $0 \leq \tilde{\theta} < \pi$ and $0 \leq \tilde{\phi} < 2\pi$ and we will denote by σ the radial coordinate of the CY cone which, together with an angular coordinate ψ , will parameterize the normal bundle N_2 . Moreover, we shall choose a system of spherical coordinates for the transverse \mathbb{R}^3 , r being the corresponding radial coordinate and (θ, ϕ) the angular variables ($0 \leq \theta < \pi$, $0 \leq \phi < 2\pi$). The concrete ansatz for the ten-dimensional string frame metric we will adopt is the following:

$$\begin{aligned}
 ds_{st}^2 = & e^{2\Phi} \left[dx_{1,2}^2 + \mathcal{Z} R^2 \left(d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2 \right) \right] + \\
 & + e^{-2\Phi} \left[\frac{1}{\mathcal{Z}} \left(d\sigma^2 + \sigma^2 \left(d\psi + \cos \tilde{\theta} d\tilde{\phi} \right)^2 \right) + dr^2 + r^2 \left(d\theta^2 + \sin^2 \theta d\phi^2 \right) \right], \quad (2.1)
 \end{aligned}$$

where $dx_{1,2}^2$ denotes the Minkowski metric in 2+1 dimensions and the range of ψ is $0 \leq \psi < 2\pi$. Notice that ψ is fibered over the (θ, ϕ) two-sphere. For convenience we have included in (2.1) the radius R , given by:

$$R^3 = 8\pi g_s N_c (\alpha')^{\frac{3}{2}}, \quad (2.2)$$

with g_s and α' being respectively the string coupling constant and the Regge slope. The ansatz (2.1) contains two functions: \mathcal{Z} , which controls the size of the cycle, and Φ , which is the dilaton of the type IIA theory. Both of them should be considered as functions of the two radial coordinates r and σ :

$$\Phi = \Phi(r, \sigma), \quad \mathcal{Z} = \mathcal{Z}(r, \sigma). \quad (2.3)$$

As in any other background generated by D4-branes, the ansatz should be endowed with an RR four-form F_4 . Let C_3 denote the corresponding three-form potential ($F_4 = dC_3$). We shall adopt the following ansatz for C_3 :

$$C_3 = -g \omega_2 \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}), \quad (2.4)$$

where $g(r, \sigma)$ is a new function and ω_2 is the volume element of the (θ, ϕ) two-sphere:

$$\omega_2 = \sin \theta d\theta \wedge d\phi. \quad (2.5)$$

The corresponding RR four-form field strength will be:

$$F_4 = -(\dot{g} d\sigma + g' dr) \wedge \omega_2 \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) + g \tilde{\omega}_2 \wedge \omega_2, \quad (2.6)$$

where we have denoted:

$$' \equiv \partial_r, \quad \dot{} \equiv \partial_\sigma. \quad (2.7)$$

We will require that our background preserves eight supersymmetries. which is the appropriate number of SUSYs for a supergravity dual of $\mathcal{N} = 4$ gauge theories in three dimensions. As shown in detail in appendix A, the vanishing of the different components of the supersymmetric variations of the gravitino and dilatino gives rise to a system of first-order BPS equations for the functions Φ , \mathcal{Z} and g entering our ansatz. This system is the following:

$$\begin{aligned} g &= -R^2 r^2 \mathcal{Z}', & e^{-4\Phi} \sigma &= R^2 \mathcal{Z} \dot{\mathcal{Z}} \\ g' &= -4\sigma r^2 e^{-4\Phi} \dot{\Phi}, & \dot{g} &= -\sigma R^{-2} \mathcal{Z}^{-2} e^{-4\Phi} g + 4\sigma r^2 \mathcal{Z}^{-1} e^{-4\Phi} \Phi'. \end{aligned} \quad (2.8)$$

It is interesting to notice that not all the equations in (2.8) are independent. Actually, one can check that the equation for \dot{g} in (2.8) can be obtained from the others. Moreover, one can combine the different equations in (2.8) and get a single second-order PDE for the function $\mathcal{Z}(r, \sigma)$, namely:

$$r \mathcal{Z} \left(\dot{\mathcal{Z}} - \sigma \ddot{\mathcal{Z}} \right) = \sigma \left(r \dot{\mathcal{Z}}^2 + r \mathcal{Z}'' + 2\mathcal{Z}' \right). \quad (2.9)$$

Notice that, if \mathcal{Z} is known, the other functions Φ and g can be determined from the first two equations in (2.8). Moreover, we check in appendix A that the second order equations of motion for the RR four-form F_4 , dilaton Φ and the metric G_{MN} of type IIA supergravity follow from the system (2.8).

2.1 Integration of the BPS system

The BPS system (2.8) can be integrated by elementary methods when $\sigma = 0$ and r varies. Indeed, it follows from the last line in (2.8) that $g(r, \sigma = 0)$ is constant. Let us put:

$$g(r, \sigma = 0) = g_0. \quad (2.10)$$

Then, the first equation in (2.8) for $\sigma = 0$ can be readily integrated, namely:

$$\mathcal{Z}'(r, 0) = -\frac{g_0}{r^2 R^2} \Rightarrow \mathcal{Z}(r, 0) = \frac{g_0}{r R^2} + \text{constant}. \quad (2.11)$$

The actual value of g_0 can be obtained from the quantization condition of the RR four-form flux F_4 . Actually, the best way to perform this analysis is by using the approach in which the solution is obtained by uplifting from gauged supergravity. This study was done in refs. [9, 10] and allows one to find a solution of the system (2.8) for arbitrary values of the

variables r and σ . Here we will just reproduce this solution with our notations. First of all, let us define the function $\Gamma(\mathcal{Z})$ as follows:

$$\Gamma(\mathcal{Z}) \equiv \mathcal{Z}_* + 2(\mathcal{Z} - \mathcal{Z}_*) + \frac{16\kappa}{\mathcal{Z}_*} (\mathcal{Z} - \mathcal{Z}_*)^2, \quad (2.12)$$

where \mathcal{Z}_* and κ are constants. Then, $\mathcal{Z}(r, \sigma)$ is determined implicitly as the solution of the equation:

$$(\mathcal{Z}_* - \mathcal{Z}) \left[r^2 + \frac{\sigma^2}{\Gamma(\mathcal{Z})} \right]^{\frac{1}{2}} = \frac{R}{8}. \quad (2.13)$$

We are able to solve this quartic equation in \mathcal{Z} exactly, although we must be careful that when we use this solution in the following we are always picking up the appropriate root. For each value of r and σ it must be checked that the solution is real and less than \mathcal{Z}_* . Notice from (2.13) that \mathcal{Z}_* is just the constant value approached by \mathcal{Z} in the UV, i.e. when either r or σ is large. Moreover, $g(r, \sigma)$ and $\Phi(r, \sigma)$ are given by:

$$g = \frac{R^2 (\mathcal{Z} - \mathcal{Z}_*) r^3}{r^2 + \frac{\mathcal{Z}}{\Gamma^2(\mathcal{Z})} \sigma^2}, \quad e^{-4\Phi} = R^2 \frac{\mathcal{Z} (\mathcal{Z}_* - \mathcal{Z})}{\left[r^2 + \frac{\mathcal{Z}}{\Gamma^2(\mathcal{Z})} \sigma^2 \right] \Gamma(\mathcal{Z})}. \quad (2.14)$$

As a check of this solution, one can easily prove by taking derivatives of (2.12)–(2.13) that the BPS system (2.8) is satisfied. Moreover, by taking $\sigma = 0$ in (2.13) one can verify that $\mathcal{Z}(r, 0)$ is indeed of the form (2.11), with g_0 being given by:

$$g_0 = -\frac{R^3}{8}. \quad (2.15)$$

Using this result we can rewrite $\mathcal{Z}(r, 0)$ as:

$$\mathcal{Z}(r, 0) = \mathcal{Z}_* - \frac{R}{8r}. \quad (2.16)$$

Furthermore, by using (2.16) to evaluate the right-hand side of the second equation in (2.14), one can obtain the value of the dilaton at $\sigma = 0$, namely:

$$e^{-4\Phi(r, 0)} = \frac{\mathcal{Z}_* - \frac{R}{8r}}{\mathcal{Z}_* - \frac{R}{4r} + \frac{\kappa}{4\mathcal{Z}_*} \frac{R^2}{r^2}} \frac{R^3}{r^3}. \quad (2.17)$$

From the explicit expressions for \mathcal{Z} and Φ written above one easily concludes that, when r and σ are small enough, the supergravity solution is not valid because the function \mathcal{Z} becomes negative and/or the dilaton Φ becomes complex. This phenomenon is related to the so-called enhancement mechanism [26] (see below). We can estimate the scale at which this mechanism occurs by computing from (2.16) the value of r for which $\mathcal{Z}(r, 0)$ vanishes. This determines the so-called enhancement radius r_e , given by:

$$r_e = \frac{R}{8\mathcal{Z}_*}. \quad (2.18)$$

Notice also that the sign of the right-hand side of (2.17) becomes negative for sufficiently small r . Actually, the numerator in (2.17) changes its sign precisely at $r = r_e$, whereas

the change of sign of the denominator depends on the value of the constant κ . Indeed, by analyzing the discriminant of the quadratic function in the denominator of (2.17) one easily concludes that this equation has no real roots for $\kappa > 1/16$. Thus in this case $e^{-4\Phi(r,0)}$ changes its sign precisely at the enhançon radius r_e . However, for $\kappa < \frac{1}{16}$ this change of sign occurs for larger values of r , namely for $r = r_H$, where r_H is given by:

$$r_H = \left(1 + \sqrt{1 - 16\kappa}\right) r_e . \tag{2.19}$$

Clearly, when $\kappa \leq 1/16$ the space ends, at $\sigma = 0$, when $r = r_H \geq r_e$.

2.2 UV form of the metric

As mentioned above, it follows from (2.13) that the function \mathcal{Z} approaches the constant value $\mathcal{Z} = \mathcal{Z}_*$ as one moves into the UV region. Since the solution of the algebraic equation (2.13) for $\mathcal{Z}(r, \sigma)$ is complicated, we can try to perform an expansion around this constant value. Keeping the first non-trivial term, $\mathcal{Z}(r, \sigma)$ can be approximated by the following expression:

$$\mathcal{Z}(r, \sigma) \approx \mathcal{Z}_* - \frac{R}{8} \frac{\mathcal{Z}_*^{1/2}}{\sqrt{r^2 \mathcal{Z}_* + \sigma^2}} , \tag{2.20}$$

while for the dilaton Φ we have:

$$e^{-4\Phi(r, \sigma)} \approx \frac{R^3}{8} \frac{\mathcal{Z}_*^{3/2}}{[\sqrt{r^2 \mathcal{Z}_* + \sigma^2}]^3} . \tag{2.21}$$

Notice that the expression for \mathcal{Z} in (2.20) gives the exact result (2.16) for $\sigma = 0$. The above analysis suggests that in the UV region the combination $r^2 \mathcal{Z}_* + \sigma^2$ plays a significant role. Having this in mind we define a new set of variables, u and $\hat{\alpha}$, as follows:

$$u = \sqrt{r^2 \mathcal{Z}_* + \sigma^2} \quad \& \quad \tan \hat{\alpha} = \frac{\sigma}{\sqrt{\mathcal{Z}_*} r} \quad \text{with} \quad 0 < \hat{\alpha} < \frac{\pi}{2} . \tag{2.22}$$

The function \mathcal{Z} as well as the dilaton generally depend on both coordinates, u and $\hat{\alpha}$, but in the UV limit of large u the $\hat{\alpha}$ dependence disappears. Actually, their expressions when $u \rightarrow \infty$ are:

$$\mathcal{Z} \rightarrow \mathcal{Z}_* \quad \& \quad e^{-2\Phi} \rightarrow \frac{\mathcal{Z}_*^{3/4}}{2\sqrt{2}} \left(\frac{R}{u}\right)^{3/2} . \tag{2.23}$$

Using these values in the metric ansatz we end up with the following expression:

$$ds_{UV}^2 \approx \frac{2\sqrt{2}}{\mathcal{Z}_*^{3/4}} \left(\frac{u}{R}\right)^{3/2} \left[dx_{1,2}^2 + \mathcal{Z}_* R^2 d\tilde{\Omega}_2^2\right] + \frac{1}{2\sqrt{2} \mathcal{Z}_*^{1/4}} \left(\frac{R}{u}\right)^{3/2} du^2 + \frac{1}{2\sqrt{2}} \frac{R^{3/2}}{\mathcal{Z}_*^{1/4}} u^{1/2} \left[d\hat{\alpha}^2 + \cos^2 \hat{\alpha} d\Omega_2^2 + \sin^2 \hat{\alpha} (d\psi + \cos \tilde{\theta} d\tilde{\phi})^2\right] , \tag{2.24}$$

where $d\tilde{\Omega}_2^2 \equiv d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2$ is the line element of the $(\tilde{\theta}, \tilde{\phi})$ two-sphere. In order to interpret the meaning of the results just found, let us recall that, given a background

of type IIA theory such as the one we are considering, one can generate a solution of eleven-dimensional supergravity by uplifting the metric by means of the standard formula:

$$ds_{11}^2 = e^{-\frac{2}{3}\Phi} ds_{10}^2 + e^{\frac{4}{3}\Phi} (dz)^2, \tag{2.25}$$

where z is the eleventh M-theory coordinate. We shall apply (2.25) to the ten-dimensional UV metric and dilaton written in eqs. (2.24) and (2.23). After changing the radial variable u by a new coordinate y , defined as:

$$y^2 = \frac{2R}{\sqrt{\mathcal{Z}_*}} u, \tag{2.26}$$

the resulting eleven-dimensional UV metric becomes:

$$ds_{11}^2 \approx \frac{y^2}{R^2} \left[dx_{1,3}^2 + \mathcal{Z}_* R^2 d\tilde{\Omega}_2^2 \right] + R^2 \left(\frac{dy}{y} \right)^2 + \frac{R^2}{4} \left[d\hat{\alpha}^2 + \cos^2 \hat{\alpha} d\Omega_2^2 + \sin^2 \hat{\alpha} (d\psi + \cos \tilde{\theta} d\tilde{\phi})^2 \right], \tag{2.27}$$

where $dx_{1,3}^2 = dx_{1,2}^2 + dz^2$. From (2.27) we conclude that the uplifted metric is of the form $AdS_7 \times S^4$, with the AdS_7 having two of its directions compactified in a two-sphere and with the S^4 being fibered over this S^2 . Notice also that the radius of the AdS_7 is just R , whereas the S^4 has radius $R/2$. These results are, of course, consistent with the origin of the solution [9], as coming from M5 wrapped on a two-cycle.

3. Addition of flavor branes

In this section we will start exploring the possibility of finding the dual of the $\mathcal{N} = 4$ 3d gauge theory with matter hypermultiplets in the fundamental representation of the gauge group. We will achieve this by adding flavor branes to the setup of section 2. These flavor branes will be D4-branes extended along the three Minkowski directions x^0, x^1 and x^2 as well as the ψ and σ directions of the Calabi-Yau. At the same time they will be located at particular fixed values of the S^2 sphere and of the transverse \mathbb{R}^3 , as represented in the following array:

	$\mathbb{R}^{1,2}$			$\overbrace{S^2}^{CY_2}$		N_2		\mathbb{R}^3		
N_c D4 (color)	—	—	—	○	○	·	·	·	·	·
N_f D4 (flavor)	—	—	—	·	·	—	—	·	·	·

It is worth pointing out that, in the above setup, the flavor branes wrap a non-compact direction of the internal Calabi-Yau. This is, actually, a standard requirement which one should demand of these kinds of setup in order to convert, in the appropriate decoupling limit, the gauge symmetry living on the worldvolume of the flavor brane into a global (flavor) symmetry.

To describe more precisely the embedding of the flavor brane let us choose the following system of worldvolume coordinates:

$$\xi^\alpha = (x^0, x^1, x^2, \sigma, \psi) . \tag{3.1}$$

Then, the embedding is determined by the condition that all other remaining ten-dimensional coordinates are constant. By analyzing the kappa symmetry of the worldvolume action of the flavor brane we will explicitly check in appendix B that the embedding just described preserves the same supersymmetry as the unflavored setup. Moreover, the position in r of the flavor brane, r_Q , represents the distance between the two sets of branes and has a well-defined meaning in the gauge theory dual. Indeed, r_Q is related to the mass m_Q of the matter hypermultiplet by means of the following relation:

$$m_Q = \frac{r_Q}{2\pi\alpha'} . \tag{3.2}$$

In the so-called quenched approximation the effects of the quark loops in the field theory observables are neglected. This approximation is well justified when $N_f \ll N_c$ and corresponds, on the gravity side, to treating the flavor branes as probes and neglecting their influence on the metric. Later in this paper we will make use of this approximation to study several aspects of the gauge theory, such as the meson spectrum and the quark-antiquark potentials. However, in the remainder of this section we will analyze, in a certain approximation, how the backreaction of the flavor branes modifies the solution described in section 2.

3.1 Including the backreaction

Let us study the backreaction of the flavor branes on the background in the case in which the number of flavors N_f is large and of the same order as the number of colors N_c . From the field theory point of view this limit was considered a long time ago by Veneziano [6]. Here we will follow the approach pioneered in ref. [11], which is based on the observation that, when $N_f \rightarrow \infty$, one can homogeneously distribute the N_f flavor branes in their transverse directions (for a clear discussion on the validity of the DBI+WZ action for a large number of smeared branes, see section 7 of [21]). Notice that, when the branes are embedded as explained around (3.1), they preserve the same supersymmetries independent of their position in the transverse space. Actually, we will consider a distribution of branes with a fixed value r_Q of the r coordinate and smeared along the angular coordinates (θ, ϕ) and $(\tilde{\theta}, \tilde{\phi})$. In order to figure out how this smearing is implemented, let us recall that the action for a stack of N_f D4-branes is given by the sum of the DBI and WZ terms:

$$S_{\text{flavor}} = -T_4 \sum_{N_f} \int_{\mathcal{M}_5} d^5\xi e^{-\Phi} \sqrt{-\det \hat{G}_5} + T_4 \sum_{N_f} \int_{\mathcal{M}_5} \hat{C}_5 , \tag{3.3}$$

with \hat{G}_5 being the induced metric on the worldvolume and \hat{C}_5 the pullback of the RR four-form potential to \mathcal{M}_5 . The smearing procedure amounts to promoting the infinite

sum appearing in the action to a ten-dimensional integral. For the WZ part we have:

$$\sum_{N_f} \int_{\mathcal{M}_5} \hat{C}_5 \rightarrow \int_{\mathcal{M}_{10}} \Omega \wedge C_5, \quad (3.4)$$

where Ω is a five-form proportional to the volume form of the transverse space, namely:

$$\Omega = \frac{N_f}{16\pi^2} \delta(r - r_Q) dr \wedge \omega_2 \wedge \tilde{\omega}_2, \quad (3.5)$$

with ω_2 being given by (2.5) and $\tilde{\omega}_2 = \sin \tilde{\theta} d\tilde{\theta} \wedge d\tilde{\phi}$. The smearing form Ω is normalized as:

$$\int \Omega = N_f. \quad (3.6)$$

Similarly, to obtain the smeared version of the DBI part of the flavor brane action we perform the substitution:

$$\sum_{N_f} \int_{\mathcal{M}_5} d^5 \xi e^{-\Phi} \sqrt{-\det \hat{G}_5} \rightarrow \int_{\mathcal{M}_{10}} d^{10} x e^{-\Phi} \sqrt{-\det G} |\Omega|, \quad (3.7)$$

where $|\Omega|$ is the modulus of Ω :

$$|\Omega| = \sqrt{\frac{1}{5!} \Omega_{M_1 \dots M_5} \Omega_{N_1 \dots N_5} \prod_{k=1}^5 G^{M_k N_k}}. \quad (3.8)$$

Therefore, the smeared DBI+WZ action of the flavor branes is:

$$S_{\text{flavor}} = -T_4 \int_{\mathcal{M}_{10}} d^{10} x e^{-\Phi} \sqrt{-\det G} |\Omega| + T_4 \int_{\mathcal{M}_{10}} \Omega \wedge C_5. \quad (3.9)$$

By inspecting the WZ term in the action (3.9) one readily concludes that the flavor brane acts as a source for the RR six-form $F_6 = dC_5$ which, in turn, induces a violation of the Bianchi identity of $F_4 = *F_6$. Actually, one can prove that this modified Bianchi identity becomes:

$$dF_4 = 2\kappa_{10}^2 T_4 \Omega = \frac{N_f R^3}{2N_c} \delta(r - r_Q) dr \wedge \omega_2 \wedge \tilde{\omega}_2, \quad (3.10)$$

where, in the last step, we have used that $2\kappa_{10}^2 = (2\pi)^7 g_s^2 (\alpha')^4$ and $T_4 = 1/(2\pi)^4 g_s (\alpha')^{\frac{5}{2}}$ and, thus, $2\kappa_{10}^2 = \pi^2 R^3/N_c$.

Let us now formulate a new ansatz for this backreacted flavored setup. First of all, we will adopt the same ansatz (2.1) for the ten-dimensional metric. However, as is clear from (3.10), we should change the ansatz for F_4 in order to reproduce the modified Bianchi identity. Actually, the natural ansatz both satisfying (3.10) and generalizing (2.6) is:

$$F_4 = -(\dot{g} d\sigma + g' dr) \wedge \omega_2 \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) + \left(g + \frac{N_f R^3}{2N_c} \Theta(r - r_Q) \right) \tilde{\omega}_2 \wedge \omega_2, \quad (3.11)$$

where Θ is the Heaviside step function. Proceeding as in the unflavored case and substituting the new ansatz (3.11) for F_4 into the equations for the supersymmetric variations of dilatino and gravitino we have:

$$\begin{aligned} \left[g + \frac{N_f R^3}{2N_c} \Theta(r - r_Q) \right] &= - R^2 r^2 \mathcal{Z}', \\ e^{-4\Phi} \sigma &= R^2 \mathcal{Z} \dot{\mathcal{Z}}, \\ g' &= -4\sigma r^2 e^{-4\Phi} \dot{\Phi}, \\ \dot{g} &= -\sigma R^{-2} \mathcal{Z}^{-2} e^{-4\Phi} \left[g + \frac{N_f R^3}{2N_c} \Theta(r - r_Q) \right] \\ &\quad + 4\sigma r^2 \mathcal{Z}^{-1} e^{-4\Phi} \Phi'. \end{aligned} \tag{3.12}$$

The set of projections on the Killing spinors needed to arrive at (3.12) is the same as in the unflavored case. Thus, the flavored solutions preserve also the same eight supersymmetries as the unflavored ones. Moreover, as in (2.8), the last equation in the system (3.12) is not independent of the others. One can also verify (see appendix A) that the equations of motion of F_4 and the metric (including the contribution of the DBI action to Einstein's equations) are satisfied if the system (3.12) holds. The analogue of the PDE (2.9) for the flavored case is:

$$r^2 \mathcal{Z} \left(\dot{\mathcal{Z}} - \sigma \ddot{\mathcal{Z}} \right) = r \sigma \left(r \dot{\mathcal{Z}}^2 + r \mathcal{Z}'' + 2\mathcal{Z}' \right) + \sigma \frac{N_f R^3}{2N_c} \frac{R^3}{8} \delta(r - r_Q). \tag{3.13}$$

As in the unflavored case, we can integrate the function \mathcal{Z} for $\sigma = 0$. Indeed, it follows from (3.12) that g is independent of r when the variable σ vanishes. If we call g_0 this constant value of g then, for $r > r_Q$, we have:

$$\mathcal{Z}'(r, 0) = -\frac{1}{R^2} \left[g_0 + \frac{N_f R^3}{2N_c} \frac{R^3}{8} \right] \frac{1}{r^2}, \quad (r > r_Q), \tag{3.14}$$

and after using for g_0 the same value as in the unflavored case, we get:

$$\mathcal{Z}'(r, 0) = \frac{R}{8r^2} \left[1 - \frac{N_f}{2N_c} \right], \quad (r > r_Q). \tag{3.15}$$

When $r < r_Q$, by simply putting $N_f = 0$ on the right-hand side of (3.15), we recover the unflavored result. Integrating (3.15) and imposing continuity for the solution along $r = r_Q$ we have:

$$\mathcal{Z}(r, 0) = \mathcal{Z}_* - \frac{R}{8r_Q} \frac{N_f}{2N_c} \Theta(r - r_Q) - \frac{R}{8r} \left[1 - \frac{N_f}{2N_c} \Theta(r - r_Q) \right], \tag{3.16}$$

where the constant \mathcal{Z}_* is the same as in (2.16). We have not been able to obtain analytically the solution of the flavored system (3.12) for arbitrary values of r and σ . Instead, one can integrate numerically this system of equations. Since the system (3.12) reduces to the unflavored one written in (2.8), we can assume that the solution of (3.12) reduces to the

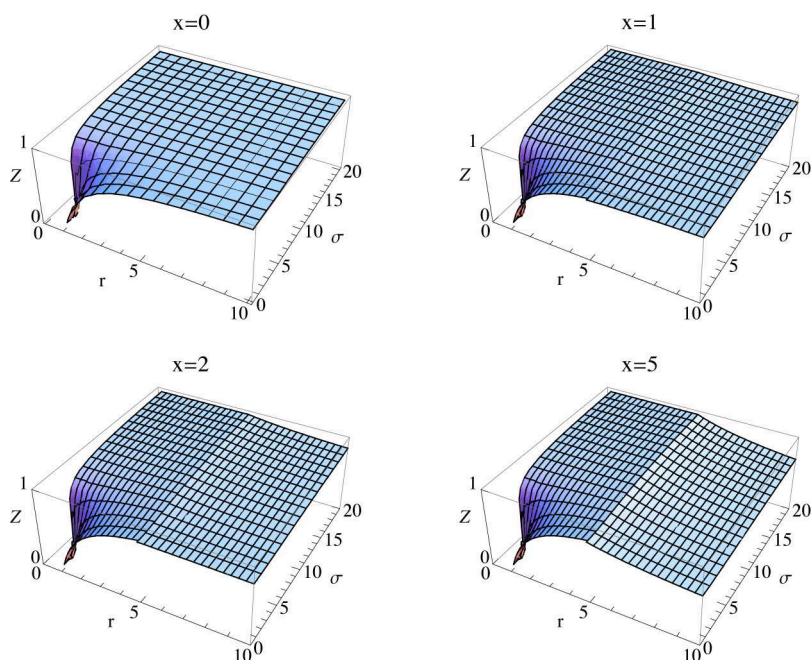


Figure 1: Solutions for Z for different values of $x = \frac{N_f}{N_c}$ as a function of r and σ . These curves correspond to $\kappa = Z_* = 1$ and $r_Q = 5$.

one written in (2.12)–(2.14) for $r < r_Q$, while at $r = r_Q$ the functions Z , g and Φ are continuous and Z' has a discontinuity independent of σ and given by:

$$Z'(r_Q + \epsilon, \sigma) - Z'(r_Q - \epsilon, \sigma) = -\frac{R}{8r_Q^2} \frac{N_f}{2N_c}. \quad (3.17)$$

Using the continuity and change in derivative of Z at $r = r_Q$ and 3.16 as the boundary conditions, we are able to numerically integrate the equation of motion for Z up to a finite value of r after which the solution becomes highly oscillatory and unstable. The result of the numerical integration for $Z(r, \sigma)$ is shown in figure 3.1 for several values of N_c and N_f . It is quite evident from figure 3.1 that Z has a wedge shape at the position $r = r_Q$ of the flavor branes. This means that there will be a curvature singularity at this point, which is actually needed to match a similar term in the Einstein equations coming from the energy-momentum tensor of the flavor brane sources. As in the unflavored case, the geometry is also singular in the IR due to the enhançon phenomenon. Outside these regions the metric is regular and one can verify that the components of the Ricci tensor depend on N_c and N_f through the combination $N_c - \frac{N_f}{2}$, which scales as N_c in the Veneziano limit $N_c, N_f \rightarrow \infty$ with N_c/N_f fixed. Thus, outside the location of the source, the region in which the supergravity approximation is valid is not modified by the backreaction of the flavor branes, as happened in the backgrounds of refs. [11, 15, 22].

4. Connection with gauge theory

The $\mathcal{N} = 4$ three-dimensional gauge theories can be regarded as the reduction to three dimensions of the $\mathcal{N} = 2$ gauge theories in four dimensions [27]. The field content of the different 3d supermultiplets can be easily obtained by dimensional reduction of the one corresponding to $\mathcal{N} = 2, d = 4$. Let us consider, for example, the vector multiplet which, in four dimensions is composed by a vector field, a Dirac spinor and a complex scalar. By reducing to three dimensions one gets one vector, two Dirac fermions and three scalars. Similarly, one can verify that the $\mathcal{N} = 2, d = 3$ hypermultiplet contains two Dirac fermions and four complex scalars. Given this field content it is straightforward to find the one loop running coupling constant $g_{\text{YM}}(\mu)$ of the theory (see, for example, the appendix C of [10]). Indeed, let us consider a $\text{SU}(N_c)$ gauge theory with N_f matter hypermultiplets in the fundamental representation. Then, one can show that:

$$\frac{1}{g_{\text{YM}}^2(\mu)} = \frac{1}{g_{\text{YM}}^2} \left[1 - \frac{g_{\text{YM}}^2 N_c}{4\pi\mu} \left(1 - \frac{N_f}{2N_c} \right) \right], \quad (4.1)$$

where μ is the energy scale. The one-loop result displayed in eq. (4.1) is, actually, exact in perturbation theory. Notice that (4.1) shows that the $\mathcal{N} = 4$ theory has negative beta function when $N_f < 2N_c$, while for $N_f > 2N_c$ the beta function changes its sign and becomes positive. In the borderline case $N_f = 2N_c$ the one-loop beta function vanishes and the coupling does not run anymore in perturbation theory. In the next subsection we will show that our gravity solutions match perfectly the behavior (4.1), both in the unflavored and backreacted flavored cases.

Besides the perturbative result just reviewed, the $\mathcal{N} = 4$ 3d theories have a very rich non-perturbative structure. Indeed, the Coulomb branch of vacua of these theories is a hyperkähler manifold, which is isomorphic to the moduli space of three-dimensional monopole solutions of a different gauge theory [27, 28]. Moreover, the Higgs branch is also a hyperkähler manifold. Furthermore, these theories display the phenomenon of mirror symmetry [29, 30], which is a duality between two different $\mathcal{N} = 4$ 3d gauge theories which exchanges the Higgs and Coulomb branches, as well as the Fayet-Iliopoulos and mass terms. When these $\mathcal{N} = 4$ theories are realized in the type IIB string theory as in [28], the mirror symmetry is just a manifestation of the underlying S-duality of the type IIB string theory (see ref. [31] for a review).

4.1 Probe calculation

In order to extract information about the gauge theory living on the D4-branes, we will study, following ref. [10], the dynamics of a color D4-brane probe wrapping an S^2 and moving under the influence of the metric and RR form of the background. The action of such a probe will be:

$$S = -T_4 \int d^5\xi e^{-\Phi} \sqrt{-\det(\hat{G}_5 + 2\pi\alpha' F)} + T_4 \int (\hat{C}_5 + 2\pi\alpha' \hat{C}_3 \wedge F), \quad (4.2)$$

where ξ^a ($a = 0, \dots, 4$) is the set of worldvolume coordinates along which the color D4-brane is extended, F is the field strength for the worldvolume gauge field and the hat over

G_5 , C_5 and C_3 denotes the pullback over the worldvolume of the D4-brane. In particular, let us evaluate the action (4.2) in the case in which the ten-dimensional metric is of the form (2.1). Let us choose the worldvolume coordinates as $\xi^a = (x^0, x^1, x^2, \tilde{\theta}, \tilde{\phi})$ and let us assume that the brane is embedded in such a way that the remaining ten-dimensional coordinates are constant and the worldvolume gauge field F is zero. Then, the determinant of the induced metric for such a configuration takes the form:

$$e^{-\Phi} \sqrt{-\det \hat{G}_5} = R^2 \mathcal{Z} e^{4\Phi} \sin \tilde{\theta} \sqrt{1 + \sigma^2 \frac{\cot^2 \tilde{\theta}}{R^2 \mathcal{Z}^2 e^{4\Phi}}} . \quad (4.3)$$

For the configuration we are considering, the only non-vanishing contribution to the WZ part in (4.2) is the term containing C_5 . The expression of this five-form potential for our solutions has been evaluated in appendix A (eq. (A.9)). The corresponding pullback to the D4-brane worldvolume is:

$$\hat{C}_5 = R^2 \mathcal{Z} e^{4\Phi} \sin \tilde{\theta} dx^0 \wedge dx^1 \wedge dx^2 \wedge d\tilde{\theta} \wedge d\tilde{\phi} . \quad (4.4)$$

Let us now substitute (4.3) and (4.4) into (4.2). The result is just minus the static potential between the stack of N_c color branes and the probe, namely:

$$S_{\text{pot}} = -T_4 \int d^3x d\tilde{\theta} d\tilde{\phi} R^2 \mathcal{Z} e^{4\Phi} \sin \tilde{\theta} \left[\sqrt{1 + \sigma^2 \frac{\cot^2 \tilde{\theta}}{R^2 \mathcal{Z}^2 e^{4\Phi}}} - 1 \right] . \quad (4.5)$$

Notice that the right-hand side of (4.5) only vanishes when $\sigma = 0$, which should be interpreted as the point of the Calabi-Yau in which one can place a color brane without breaking supersymmetry. Actually, one can check this statement directly by studying the implementation of kappa symmetry for the different D4-brane embeddings (see subsection 4.2 and appendix B). Let us thus assume that our probe brane is located at $\sigma = 0$ and that we switch on a worldvolume gauge field $F_{\mu\nu}$ whose only non-vanishing components are those along the Minkowski directions x^μ . We will expand the DBI lagrangian density (integrated over the angular directions $\tilde{\theta}$ and $\tilde{\phi}$) up to quadratic order in the gauge field $F_{\mu\nu}$. The result of this expansion can be parameterized as:

$$\int d\tilde{\theta} d\tilde{\phi} \mathcal{L}_{\text{DBI}} \Big|_{\text{quadratic}} = -\frac{1}{2g_{\text{YM}}^2(\mu)} \text{tr}[F_{\mu\nu} F^{\mu\nu}] , \quad (4.6)$$

where $g_{\text{YM}}(\mu)$ is, by definition, the Yang-Mills coupling at the renormalization scale μ . Actually, by performing explicitly the calculation, one gets:

$$\frac{1}{g_{\text{YM}}^2(\mu)} = \frac{R^2}{2\pi g_s (\alpha')^{1/2}} \mathcal{Z}(r, \sigma = 0) . \quad (4.7)$$

Clearly, the variable r should be related to the energy scale μ of the field theory. The natural radius-energy relation is given by:

$$r = 2\pi \alpha' \mu , \quad (4.8)$$

which we will assume to hold in the following. Let us now suppose that $r > r_Q$ or, equivalently, that $\mu > m_Q$, where m_Q has been defined in (3.2). After substituting in (4.7) the value of $\mathcal{Z}(r, \sigma = 0)$ as given by (3.16) for the general flavored solution we have:

$$\frac{1}{g_{\text{YM}}^2(\mu)} = \frac{\mathcal{Z}_* R^2}{2\pi g_s (\alpha')^{1/2}} - \frac{N_f}{8\pi m_Q} - \frac{N_c}{4\pi\mu} \left(1 - \frac{N_f}{2N_c}\right). \quad (4.9)$$

Moreover, by defining the bare UV YM coupling as:

$$\frac{1}{g_{\text{YM}}^2} = \frac{\mathcal{Z}_* R^2}{2\pi g_s (\alpha')^{1/2}} - \frac{N_f}{8\pi m_Q}, \quad (4.10)$$

we get that eq. (4.9) matches perfectly the field theory expression (4.1). Moreover, by allowing the brane probe to move in the transverse flat space, and by looking at the action of the transverse scalar fields obtained by expanding the DBI+WZ action, one can obtain the metric of the moduli space in the Coulomb branch. After dualizing the worldvolume gauge field one can check that the moduli space is indeed hyperkähler and, actually, its metric has the form of the Taub-NUT space [10].

4.2 Higgs branch

The Higgs branch of the $\mathcal{N} = 4$ gauge theory is a phase in which the quark hypermultiplets acquire a non-vanishing expectation value. On the field theory side one can study the theory in the Higgs branch by turning on an extra Fayet-Ilioupoulos coupling in the lagrangian. The realization of this mechanism in a brane setup is well-known [28] (see [31] for a review). Indeed, as argued in ref. [28], one should reconnect the color and flavor branes in a supersymmetric way. In our holographic setup we should look for D4-brane embeddings which are compatible with all the supersymmetries of the gravity solution and such that they can be interpreted as representing a recombination of color and flavor branes (see [22, 32, 33] for a similar analysis in other brane setups). Recall that both types of D4-branes are extended along different directions of the Calabi-Yau cone. Indeed, the color branes are extended along $(\tilde{\theta}, \tilde{\phi})$ at $\sigma = 0$ while the flavor branes extend along (σ, ψ) at fixed angles $(\tilde{\theta}, \tilde{\phi})$. In order to find a configuration interpolating between these two situations it is natural to use the same system of worldvolume coordinates as in (3.1) and look for an embedding such that $\tilde{\theta}$ and $\tilde{\phi}$ are no longer constant. Instead, they will depend on the other coordinates of the CY_2 , namely:

$$\tilde{\theta} = \tilde{\theta}(\sigma, \psi), \quad \tilde{\phi} = \tilde{\phi}(\sigma, \psi). \quad (4.11)$$

In order to determine the D4-brane embeddings of the form (4.11) which preserve the supersymmetries of the background, one has to study the kappa symmetry of the brane probe. This analysis is performed in detail in appendix B. The final result found in this appendix can be nicely recast in terms of the following two complex coordinates of the CY_2 :

$$\zeta_1 \equiv \tan\left(\frac{\tilde{\theta}}{2}\right) e^{i\tilde{\phi}}, \quad \zeta_2 \equiv \sigma \sin\tilde{\theta} e^{-i\psi}. \quad (4.12)$$

It turns out that any holomorphic embedding of the type $\zeta_1 = f(\zeta_2)$, with f arbitrary, solves the kappa symmetry equations and, thus, preserves the supersymmetry of the background. In order to make contact with the field theory analysis, it is rather natural to restrict ourselves to embeddings characterized by a polynomial equation of the type:

$$\zeta_1^{p_1} \zeta_2^{p_2} = C , \tag{4.13}$$

with C being a (complex) constant and the exponents p_1 and p_2 are constant integers. Notice that, in the Coulomb branch, the embedding of the color branes correspond to eq. (4.13) with $p_1 = 0$, $p_2 = 1$ (and $C = 0$), while that of the flavor brane is given by $p_1 = 1$, $p_2 = 0$. It is thus natural to think that the Higgs branch embedding we are seeking is the one obtained by taking $p_1 = 1$, $p_2 = 1$ in (4.13). Notice that, in this case, the functions (4.11) are just given by:

$$\sin^2 \left(\frac{\tilde{\theta}}{2} \right) = \frac{\sigma_*}{\sigma} , \quad \tilde{\phi} = \psi + \tilde{\psi}_0 , \tag{4.14}$$

where σ_* and $\tilde{\psi}_0$ are constants (related to C in (4.13)). Notice that σ_* is just the minimal value of the coordinate σ (which occurs for $\tilde{\theta} = \pi$) which should correspond, in the field theory side, to the Higgs VEV.

5. Meson spectrum

In this section we analyze the mass spectrum of the mesonic excitations for the backgrounds of sections 2 and 3. In order to address this question we will add a flavor D4-brane probe to these backgrounds and we will study the normalizable fluctuations of its worldvolume fields. This analysis will lead us to a problem of the Sturm-Liouville type with a discrete set of eigenfunctions and eigenvalues.

Let us consider a flavor D4-brane in a geometry of the form (2.1). We will choose the system (3.1) of worldvolume coordinates and we will consider an embedding in which the radial coordinate r and the angular coordinates θ , ϕ , $\tilde{\theta}$ and $\tilde{\phi}$ are constants. Let us denote by r_q the constant value of r for this configuration (r_q is related to the mass m_q of the external quarks as in (3.2), i.e. $m_q = r_q/2\pi\alpha'$). The induced metric for such a configuration is:

$$\mathcal{G}_{ab} d\xi^a d\xi^b = e^{2\Phi(r_q, \sigma)} dx_{1,2}^2 + \frac{e^{-2\Phi(r_q, \sigma)}}{\mathcal{Z}(r_q, \sigma)} \left[d\sigma^2 + \sigma^2 (d\psi)^2 \right] . \tag{5.1}$$

Let us now perturb this static configuration by deforming it in the radial direction as:

$$r = r_q + \hat{r}(x^\mu, \sigma, \psi) , \tag{5.2}$$

where the fluctuation \hat{r} is small. In all the calculations of this section we will take r_q such that the $r = r_q$ surface does not enter the enhançon region and, thus, $\sigma = 0$ is the minimal value of the coordinate σ in the worldvolume. This can always be achieved by taking r_q sufficiently large.

Of course, the perturbation (5.2) is not the most general one. However, one can check that, at quadratic order, \hat{r} does not mix with other fluctuations and, therefore, it can be

studied separately. By expanding the corresponding DBI+WZ action ¹ one can verify that, up to quadratic terms, one gets that the lagrangian density for \hat{r} is:

$$\mathcal{L} = -\frac{T_4}{2} e^{-3\Phi} \sqrt{-\det \mathcal{G}} \mathcal{G}^{ab} \partial_a \hat{r} \partial_b \hat{r} = -\frac{T_4}{2} \left[\frac{\sigma}{e^{4\Phi} \mathcal{Z}} (\partial_{x^\mu} \hat{r})^2 + \sigma (\partial_\sigma \hat{r})^2 + \frac{1}{\sigma} (\partial_\psi \hat{r})^2 \right], \quad (5.3)$$

where Φ and \mathcal{Z} should be understood as functions of σ at $r = r_q$. The equation derived from the lagrangian (5.3) is given by:

$$\partial_\sigma [\sigma \partial_\sigma \hat{r}] + \frac{\sigma}{e^{4\Phi} \mathcal{Z}} \partial_{x^\mu}^2 \hat{r} + \frac{1}{\sigma} \partial_\psi^2 \hat{r} = 0. \quad (5.4)$$

To find the solutions of this equation, let us separate variables as:

$$\hat{r} = \chi(\sigma) e^{ikx} e^{il\psi}, \quad (5.5)$$

where l is an integer (which can be taken to be non-negative without loss of generality) and k is a momentum along the Minkowski directions x^μ . Let us also define M^2 as $M^2 \equiv -k^2$, where the square is computed with the flat Minkowski metric in 2 + 1 dimensions. Plugging (5.5) into (5.4), we arrive at the following equation for $\chi(\sigma)$:

$$\partial_\sigma [\sigma \partial_\sigma \chi] + \left[\frac{\sigma}{e^{4\Phi} \mathcal{Z}} M^2 - \frac{l^2}{\sigma} \right] \chi = 0. \quad (5.6)$$

Interestingly, by means of a suitable change of variables, the fluctuation equation (5.6) can be written as a Schrödinger equation. Indeed, let us define the variable y as follows:

$$e^y = \sigma. \quad (5.7)$$

Notice that $y \in (-\infty, +\infty)$. In terms of y , the equation (5.6) of the fluctuations can be written as:

$$\frac{d^2 \chi}{dy^2} - V(y) \chi = 0, \quad (5.8)$$

where the potential $V(y)$ is given by:

$$V(y) = l^2 - M^2 \frac{e^{2y}}{e^{4\Phi} \mathcal{Z}}. \quad (5.9)$$

From the reformulation of eqs. (5.8) and (5.9) of the fluctuation equation, one can easily obtain the asymptotic value of χ when $y \rightarrow \pm\infty$. Indeed, from the behavior of $\Phi(r_q, \sigma)$ and $\mathcal{Z}(r_q, \sigma)$ when $\sigma \rightarrow 0, \infty$ one easily gets that $V \rightarrow l^2$ when $y \rightarrow \pm\infty$. It follows that, in these asymptotic regions, the two independent solutions of (5.8) when $l \neq 0$ are just $\chi \sim e^{\pm ly} = \sigma^{\pm l}$ (when $l = 0$ these solutions behave as $\chi = \text{constant}, \log \sigma$ when $\sigma \rightarrow 0, \infty$). The normalizable solutions, which can be identified with mesonic excitations, are those that are regular as $\sigma = 0, \infty$. These solutions only exist for some discrete set of values

¹One should take into account that, due to the different signs in the two projections in (A.4), the WZ coupling of the RR potential C_5 for the flavor brane must be opposite to the one appearing in (4.2).

of the mass M , which can be determined numerically by means of the shooting technique (see below).

The previous analysis only applies to the particular fluctuation \hat{r} in the radial direction of the transverse \mathbb{R}^3 . However, one can check that the same equation (5.6) describes the fluctuations of the other two \mathbb{R}^3 coordinates $\tilde{\theta}$ and $\tilde{\phi}$. Furthermore, in order to study the fluctuations of the coordinates $\tilde{\theta}$ and $\tilde{\phi}$ that determine the position of the flavor brane in the CY_2 , let us write:

$$\tilde{\theta} = \tilde{\theta}_0 + \hat{\theta}(x^\mu, \psi, \sigma), \quad \tilde{\phi} = \tilde{\phi}_0 + \hat{\phi}(x^\mu, \psi, \sigma), \quad (5.10)$$

where $\tilde{\theta}_0$ and $\tilde{\phi}_0$ are the constant unperturbed values of the angles $\tilde{\theta}$ and $\tilde{\phi}$. By plugging (5.10) into the DBI+WZ action and expanding the result up to quadratic order in the fluctuations one gets a lagrangian in which $\hat{\theta}$ and $\hat{\phi}$ are coupled. It turns out that, as happened in refs. [15] and [22], they can be easily decoupled by defining new fields $\chi_+(\sigma)$ and $\chi_-(\sigma)$ as follows:

$$\begin{aligned} \hat{\theta} &= \frac{1}{2} \left(\chi_+(\sigma) + \chi_-(\sigma) \right) \sin \theta_0 e^{ikx} \sin(l\psi), \\ \hat{\phi} &= \frac{1}{2} \left(\chi_+(\sigma) - \chi_-(\sigma) \right) e^{ikx} \cos(l\psi), \end{aligned} \quad (5.11)$$

where l is a non-negative integer. One can verify that, indeed, $\chi_+(\sigma)$ and $\chi_-(\sigma)$ satisfy two different decoupled equations. Actually, it is possible to perform a further redefinition to the functions χ_\pm in such a way that the new functions satisfy the scalar fluctuation equation (5.6). Similarly, one can show, as in ref. [15], that the fluctuations of the worldvolume gauge field also satisfy (5.6). These facts imply that, up to finite shifts in the quantum numbers, all fluctuation equations lead to the same set of eigenvalues, as expected from supersymmetry. Therefore, one can concentrate on studying the basic equation (5.6). This is what we will do in the next two subsections.

5.1 Quenched mesons in the Coulomb branch

Let us specialize the previous analysis to the case in which the background geometry is the unflavored one of section 2 and we add a flavor D4-brane probe. In this case the function \mathcal{Z} and the dilaton Φ are given by eqs. (2.13) and (2.14) respectively. The numerical results obtained by means of the shooting technique are shown in figure 2. In general, for given r_q and l , one obtains a tower of discrete normalizable states labelled by a principal quantum number n . It is interesting to notice that, for fixed n and l , the masses grow with r_q as $M^2 \sim r_q$.

One can get a rather accurate estimate of the mass spectrum by applying the semi-classical WKB method to the Schrödinger problem (5.8) (see refs. [34]–[37]). By applying this method to our particular case, we get the following mass formula:

$$M_{\text{WKB}}^2 = \frac{\pi^2}{[\zeta(r_q)]^2} (n+1)(n+3l), \quad (5.12)$$

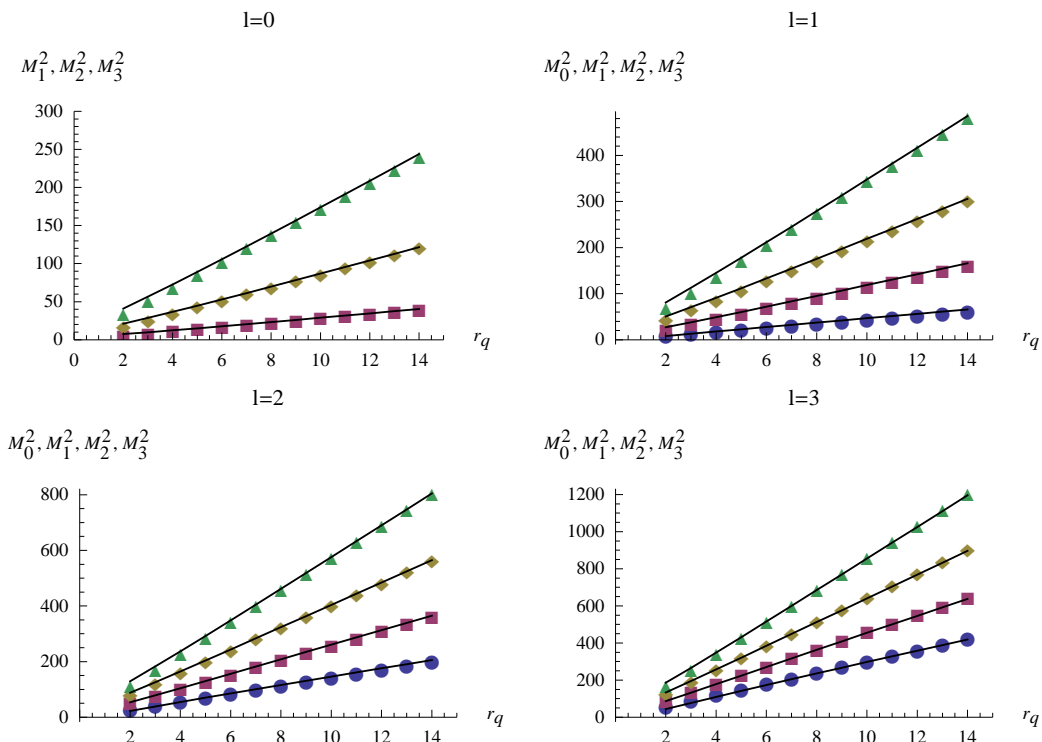


Figure 2: Mass eigenvalues from the improved WKB formula (eq. 5.16) (dots) and from the numerical calculation (solid line). The numerical calculation has been performed with the unflavored background with $\kappa = \mathcal{Z}_* = 1$.

where n is a non-negative integer ($n \geq 0$ when $l \neq 0$ and $n \geq 1$ for $l = 0$). In eq. (5.12) $\zeta(r_q)$ is the following integral:

$$\zeta(r_q) = \int_0^{+\infty} \frac{d\sigma}{\sqrt{e^{4\Phi(r_q, \sigma)} \mathcal{Z}(r_q, \sigma)}}. \quad (5.13)$$

Notice that $\zeta(r_q)$ determines the mass gap of the mesonic spectrum. One can get an approximate expression for $\zeta(r_q)$ by using inside the integral on the right-hand side of (5.13) our asymptotic expressions (2.20) and (2.21). By doing so one arrives at the following analytic expression for $\zeta(r_q)$:

$$\zeta(r_q) \approx \sqrt{\frac{R^3}{8}} \mathcal{Z}_*^{\frac{1}{4}} \int_0^{+\infty} \frac{d\sigma}{(\sigma^2 + \mathcal{Z}_* r_q^2)^{\frac{3}{4}}} = \sqrt{\frac{\pi^3 R^3}{16}} \frac{1}{[\Gamma(\frac{3}{4})]^2} \frac{1}{\sqrt{r_q}}. \quad (5.14)$$

It is interesting to stress that, within this approximation, $\zeta(r_q)$ is independent of \mathcal{Z}_* . Actually, by using the value of R written in (2.2), as well as the relation between r_q and the quark mass m_q (see (3.2)), one can write the following WKB formula for the meson masses:

$$M_{\text{WKB}}^2 = \frac{4 [\Gamma(\frac{3}{4})]^4 m_q}{\pi g_s N_c \sqrt{\alpha'}} (n+1) (n+3l). \quad (5.15)$$

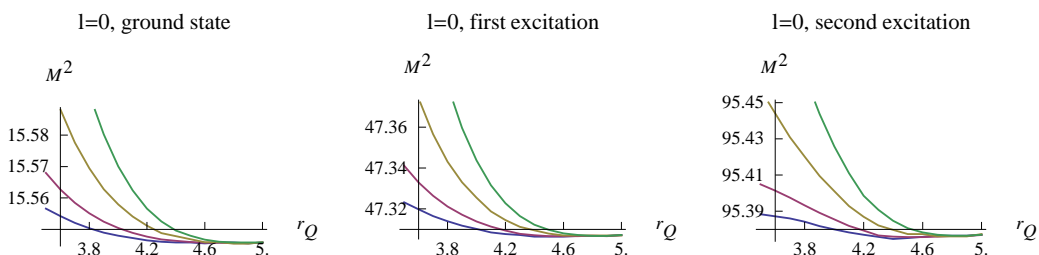


Figure 3: Variation of the meson spectrum calculated on a probe brane at $r_q = 5$ with backreacting flavor branes at varying positions r_Q . The four lines on each plot correspond to $N_f/N_c = 1$ (flattest line), $N_f/N_c = 2$, $N_f/N_c = 5$ and $N_f/N_c = 10$ (least flat line). It is clear to see that each of these return to the quenched case in the case where the backreacting branes lie close to the probe brane. All calculations have been done for $\kappa = \mathcal{Z}_* = 1$.

This formula reproduces rather well the numerical values of the masses for $l = 0, 1$ and, actually, captures accurately the meson mass gap and its dependence on the quark mass m_q . However, the degeneracies observed in the numerical results when one varies the quantum numbers n and l are not reproduced by (5.15). Nevertheless we found that the numerical results are fully recovered if (5.15) is changed to:

$$M^2 = \frac{4 \left[\Gamma\left(\frac{3}{4}\right) \right]^4 m_q}{\pi g_s N_c \sqrt{\alpha'}} (n+l) (n+2l+1). \quad (5.16)$$

The comparison between the predictions of (5.16) and the numerical values of the masses is shown in figure 2. It follows from the inspection of this figure that (5.16) provides a very good fit of the masses found by the shooting technique.

5.2 Unquenched mesons in the Coulomb branch

We can perform the analysis of the meson spectrum by looking at the fluctuations of a flavor brane probe in the background given by the backreacted flavor branes. This is equivalent to studying the fluctuations of a single quenched flavor in the presence of N_f unquenched flavors. In the large N_f limit this should be an accurate approximation. What is interesting to note here is that because the functions \mathcal{Z} and Φ are themselves continuous in r and are being taken at the value $r = r_q$, the eigenvalues for a flavor probe at this position will be exactly the same as that for a probe without the backreacted flavor.

Thus, by inspection we see from the form of the equations of motion for the probe brane fluctuations that the spectrum will be identical in the flavored and unflavored case if the probe brane is placed at the same position as the backreacting branes. The difference between the quenched and unquenched case will only be felt if we introduce a finite distance between the probe and backreacting branes, corresponding to Higgsing a $U(1)$ of the $U(N_f)$ flavor group.

The case where the probe and backreacting branes are separated is considerably more complicated, computationally, because in order to calculate the meson spectrum it is vital

to have a numerically stable geometry. The calculation of the function \mathcal{Z} from the partial differential equation is inherently an unstable calculation and therefore the calculation of the spectrum is difficult. We have however managed to calculate the spectrum in a narrow range of parameter space where the probe and backreacting branes are not very far apart.

In figure 3 we study the effects of placing the backreacting branes at a variable position, given by r_Q while always keeping the position of the probe brane at $r_q = 5$. On each graph the lines correspond to $N_f/N_c = 10, N_f/N_c = 5, N_f/N_c = 2$ and $N_f/N_c = 1$. Clearly the $N_f/N_c = 1$ line is the flattest in each case and the $N_f/N_c = 10$ line has the most variation, as expected as the geometry is most altered with a higher ratio of flavor branes to color branes. We have checked that the curves in figure 3 can be fitted to an expression of the type $M^2 = a + N_f b \left(\frac{r_Q}{r_q} - 1\right)^4$, where a and b are coefficients that are independent of N_f .

5.3 Mesons in the Higgs branch

We will now consider the fluctuations around a non-trivial embedding of the type studied in subsection 4.2 (eq. (4.13)). Recall that we argued in this subsection that these embeddings correspond to configuration in which both types of D4-branes (color and flavor) are recombined, realizing the Higgs branch in our brane setup. More concretely, we will concentrate on studying the embedding (4.13) for $p_1 = p_2 = 1$, which can be rewritten as in (4.14). We will continue to use (3.1) as our system of worldvolume coordinates but now $\tilde{\theta}$ and $\tilde{\phi}$ will no longer be constant but given by the functions displayed in (4.14). For simplicity we will restrict ourselves to studying the fluctuations of the radial coordinate r around a fixed value r_q (see eq. (5.2)). After separating variables as in (5.5) the fluctuations equation can be written in the Schrödinger form (5.8), if we introduce a new variable y as:

$$e^y = \sigma - \sigma_* , \quad -\infty < y < +\infty , \quad (5.17)$$

where σ_* is the minimal value of the coordinate σ for the embedding (see eq. (4.14)). Recall that σ_* parameterizes the Higgs VEV. The Schrödinger potential V is now given by:

$$V = \frac{l^2}{4} - M^2 \left[\frac{e^{2y}}{e^{4\Phi} \mathcal{Z}} + \sigma_* R^2 \frac{e^y \mathcal{Z}}{(e^y + \sigma_*)^2} \right]. \quad (5.18)$$

Note here that in the limit $l \rightarrow 2l$ and $\sigma_* \rightarrow 0$ we recover the Coulomb branch potential. However, one can check that this limit is not continuous and the last term in (5.18) means that the Coulomb spectrum is never recovered taking this limit in the Higgs branch.

In order to have an idea of the mass spectrum associated to the potential (5.18), let us apply the WKB method. The mass gap is now determined by an integral similar to (5.13), which is now given by:

$$\zeta(r_q, \sigma_*) = \int_{\sigma_*}^{\infty} d\sigma \sqrt{\frac{1}{e^{4\Phi(r_q, \sigma)} \mathcal{Z}(r_q, \sigma)} + \sigma_* R^2 \frac{\mathcal{Z}(r_q, \sigma)}{\sigma^2 (\sigma - \sigma_*)}}. \quad (5.19)$$

When σ_* is not very small the second term inside the square root in (5.19) dominates the integral. Actually, by applying the same approximations that led to (5.14), one arrives at the following approximate expression for $\zeta(r_q, \sigma_*)$:

$$\zeta(r_q, \sigma_*) \approx \pi R \sqrt{\mathcal{Z}_*}. \quad (5.20)$$

Notice that ζ in (5.20) is independent of r_q and σ_* and only depends on the asymptotic value of \mathcal{Z}_* . Actually, after improving the WKB formula as in (5.16), we arrive at the approximate formula for the mass levels in the Higgs branch:

$$M_{\text{Higgs}}^2 \approx \frac{1}{R^2 \mathcal{Z}_*} (n+l)(n+l+1), \quad (5.21)$$

which reproduces rather well the numerical values.

6. Wilson loops and the $q\bar{q}$ potential

In order to investigate the unflavored background more, we now turn to another direction of a more phenomenological nature, the Wilson loop operator. On the gravity side, the Wilson loop expectation value is calculated by minimizing the Nambu-Goto action for a fundamental string stretching into the dual supergravity background, whose endpoints are constrained to lie on the two sides of the Wilson loop. Below, we first briefly review the procedure for calculating Wilson loops in supergravity. We will discover that the qualitative behavior of this observable critically depends on whether the constant κ , appearing in the function (2.12), is smaller or larger than $1/16$. We will discuss these two cases separately in two different subsections.

6.1 General formalism

As stated above, the calculation of a Wilson loop in the gravity approach amounts to extremizing the Nambu-Goto action for a string propagating in the dual geometry whose endpoints trace the loop. To describe such configuration in our setup let us choose the time t and a Minkowski coordinate x as worldvolume coordinates of the string and let us consider the trajectory with:

$$r = r(x), \quad \sigma = 0, \quad \text{rest} = \text{constant}. \quad (6.1)$$

Then, the induced metric on the string worldvolume can be easily found from (2.1), namely:

$$e^{2\Phi(r,0)} \left[-dt^2 + (1 + e^{-4\Phi(r,0)} r'^2) dx^2 \right], \quad (6.2)$$

with $\Phi(r,0)$ being given by (2.17), which we can rewrite in terms of the enhancement radius r_e defined in (2.18) as:

$$e^{4\Phi(r,0)} = \frac{1}{r_e R^2 \mathcal{Z}_*} \frac{r^2 (r^2 - 2rr_e + 16\kappa r_e^2)}{r - r_e}. \quad (6.3)$$

The Nambu-Goto action for the configuration (6.1) is:

$$S = \frac{1}{2\pi} \int dt dx \sqrt{e^{4\Phi(r,0)} + r'^2}, \quad (6.4)$$

where we have taken $\alpha' = 1$ and the prime denotes a derivative with respect to x . Since the action does not explicitly depend on x , the system has a first integral r_0 , which can

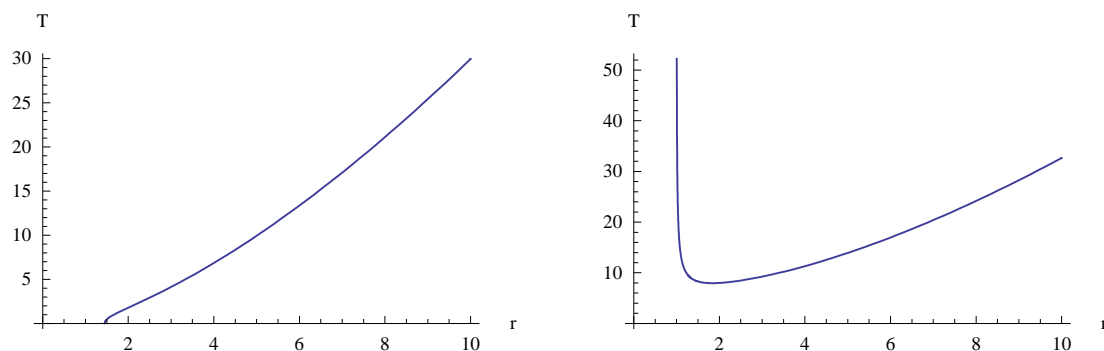


Figure 4: Effective tension of the string as a function of r . On the left is the case with $\kappa \leq 1/16$ and on the right with $\kappa > 1/16$.

be identified with the turning point of the solution. Solving the corresponding first-order equation for x in terms of r we find that the linear separation of the quark and antiquark is:

$$L = 2 \int_{r_0}^{\infty} dr \frac{e^{2\Phi(r_0,0)}}{e^{2\Phi(r,0)} \sqrt{e^{4\Phi(r,0)} - e^{4\Phi(r_0,0)}}} . \tag{6.5}$$

The energy of the configuration can be obtained from the action (6.4). Subtracting the self-energy contribution, we obtain

$$E = \frac{1}{\pi} \int_{r_0}^{\infty} dr \frac{e^{2\Phi(r_0,0)}}{\sqrt{e^{4\Phi(r,0)} - e^{4\Phi(r_0,0)}}} - \frac{1}{\pi} \int_{r_{\min}}^{\infty} dr , \tag{6.6}$$

where r_{\min} is the minimum value of r allowed by the geometry. In the specific cases, we are supposed to solve for the auxiliary parameter r_0 in terms of the separation length L . Since this cannot be done explicitly in practice, one regards (6.5) as a parametric equation for L in terms of the integration constant r_0 . Combining it with (6.6) for E , one can then determine the behavior of the potential energy of the configuration in terms of the quark-antiquark separation. Using (6.5) and (6.6) we arrive at the following relation which will prove useful in the following:

$$\frac{dE_{q\bar{q}}}{dL} = \frac{e^{2\Phi(r_0,0)}}{2\pi} . \tag{6.7}$$

The minimum value of r allowed by the geometry depends on κ . Indeed, as argued at the end of subsection 2.1, the numerator in the right-hand side of (6.3) has real roots only when $\kappa \leq 1/16$. In this case $r_{\min} = r_H$, where r_H has been written in (2.19). Conversely, when $\kappa > 1/16$ the minimum value of r is $r_{\min} = r_e$.

Following the analysis in [21], we can determine the qualitative behavior of the system simply by looking on the asymptotic expansion of the dilaton in (6.3). Utilizing their formalism we write the lagrangian as:

$$\mathcal{L} = T(r) \sqrt{1 + Y(r) r'^2} , \tag{6.8}$$

where $T(r)$ is the effective tension of the brane depending on the dilaton. In our background this is given by:

$$T(r) = \frac{1}{Y(r)^{1/2}} = e^{2\Phi(r,0)}. \quad (6.9)$$

When $\kappa \leq 1/16$ both the tension of the string and Y behave in the IR with a power law, $T = \frac{1}{Y^{1/2}} \sim (r - r_{\min})^{1/2}$. Following the reasoning of ref. [21], due to the value of the exponent of Y , the string will not reach an infinite length in the x direction when it explores the far IR of the dual theory. As we will see in the next subsection this is verified both by our numerical and analytical results.

When $\kappa > 1/16$ the string feels an effective IR wall at the point $r = r_*$ where the tension is minimized, so $T'(r_*) = 0$ and $T''(r_*) > 0$ (see figure 4). The actual value of r_* can be obtained analytically by finding the roots of a cubic polynomial. These roots depend on κ and r_e and one can check that $r_* \rightarrow 2r_e$ when κ is large. Since the background at the point $r = r_*$ is smooth and the functions of the metric do not vanish, they admit a Taylor expansion as:

$$\begin{aligned} T &= T_* + T_2(r - r_*)^2 + \dots, \quad T_2 > 0, \\ Y &= Y_* + Y_1(r - r_*) + \dots \end{aligned} \quad (6.10)$$

Following again the reasoning in [21], when the tip of the string comes close to the IR wall the length will diverge. Again this is something we will verify in the next subsection.

6.2 Case with $\kappa \leq 1/16$

For large values of r_0 the potential is, as usual, Coulombic while there exists a maximal separation $\kappa \leq 1/16$. At the values of r_0 where the separation is maximal, the energy is also maximal and always positive (see figure 5). This means that there is a screening behavior because as the potential turns positive, a configuration of two separate strings is energetically favored and corresponds to a vanishing force between the charges. As $r_0 \rightarrow r_{\min} = r_H$ we find by studying the behavior of the integrals in eqs. (6.5) and (6.6) in this case that both length and energy approach zero as:

$$L \simeq \sqrt{2} R \left(\frac{r_e Z_*}{r_H} \right)^{1/2} \sqrt{\frac{r_0}{r_H} - 1} \ln \left[\frac{3r_H}{r_0 - r_H} \right], \quad (6.11)$$

and

$$E \simeq \frac{r_0 - r_H}{2\pi} \ln \left[\frac{3r_H}{r_0 - r_H} \right]. \quad (6.12)$$

6.3 Case with $\kappa > 1/16$

In this case we follow the same steps as before but now we should have in mind that $r_{\min} = r_e$. As usual, the dependence of the potential energy for small separations L of the quark-antiquark (corresponding to large r_0) is Coulombic. In order to have an approximate expression for the separation length and the energy we have to move to the region of the

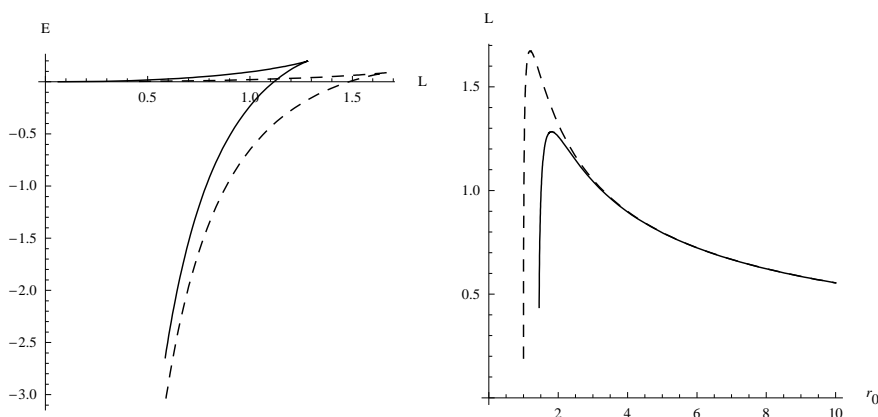


Figure 5: $\kappa \leq 1/16$: Separation length between a quark and an antiquark as a function of r_0 and energy as a function of the separation length. The dashed (solid) curve is the case for $\kappa = 1/16$ ($\kappa = 1/18$).

parameter space where κ is large. In this case the effective IR wall is located at $r = r_* \simeq 2r_e$ and the potential gives a linear confining behavior with:

$$L \simeq \frac{\mathcal{Z}_*^{1/2}}{\sqrt{3+4\kappa}} \ln \left[\frac{6r_e}{r_0 - 2r_e} \right], \quad (6.13)$$

and

$$E \simeq \frac{4r_e}{\pi} \sqrt{\frac{\kappa}{3+4\kappa}} \ln \left[\frac{6r_e}{r_0 - 2r_e} \right] \simeq \frac{4r_e}{\pi} \sqrt{\frac{\kappa}{\mathcal{Z}_*}} L. \quad (6.14)$$

For intermediate values of the separation length, the behavior of the system depends crucially on the value of κ . There is a critical value $\kappa_{\text{cr}} \simeq 63.8$, such that for $\kappa > \kappa_{\text{cr}}$ the behavior of the length and energy curves resemble the Van der Waals isotherms for a statistical system, with r_0 , L and E corresponding to volume, pressure and Gibbs potential respectively (see, for instance, [38]). In the following we will describe the situation referring to the corresponding plots in figure 6. In the region below the critical point, $\kappa < \kappa_{\text{cr}}$, the energy is a single-valued function of the length, connecting a Coulombic with a linear confining phase. For values above the critical point, $\kappa > \kappa_{\text{cr}}$, the Coulombic phase for the energy at small distances is followed by a situation where the energy becomes a triple-valued function of the length, with the physical state being the one of lowest energy. For large distances the energy returns to a single-valued function of the length with approximately linear behavior. The self-intersection point in the energy curve indicates the presence of a first-order phase transition with order parameter r_0 . At this point we should note that exactly at the critical value, $\kappa_{\text{cr}} \simeq 63.8$, the first order phase transition degenerates to a second order one. Then using purely thermodynamical arguments and the fact that not only the first but also the second derivative of the length vanishes at the critical value r_0^{cr} , we have:

$$L - L_{\text{cr}} \sim (r_0 - r_0^{\text{cr}})^3, \quad (6.15)$$

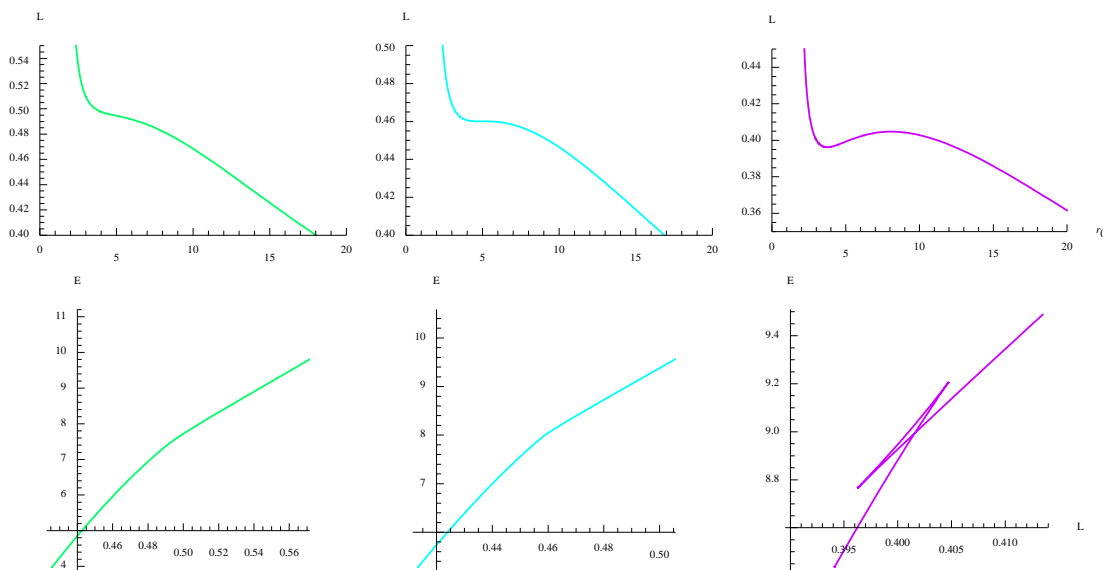


Figure 6: $\kappa > 1/16$: Separation length between a quark and an antiquark as a function of r_0 and energy as a function of the separation length. The left (right) curves correspond to the case in which κ is below (above) the critical value and the central curves are the case in which κ has exactly the critical value, $\kappa_{cr} \simeq 63.8$.

which is in agreement with gravity calculations using (6.5). The critical behavior vanishes for $\kappa < \kappa_{cr}$. When $\kappa = \kappa_{cr}$ so the corresponding critical exponent takes the classical value 3. Using (6.7) and (6.15) we can calculate the behavior of the energy close to the critical point:

$$E_{q\bar{q}} - E_{q\bar{q}}^{cr} \simeq \frac{e^{2\Phi(r_0^{cr}, 0)}}{2\pi} (L - L_{cr}) \left(1 - A|L - L_{cr}|^{1/3} \right), \quad (6.16)$$

where A is a constant. The above critical behavior is similar to that found in [39, 40, 19, 23]. This expression is calculated numerically by tuning κ very finely such that the first order phase transition, seen in $L(r_0)$, just disappears. In order to do this we must zoom into the region in the $L(r_0)$ plot where both $L'(r_0)$ and $L''(r_0)$ simultaneously vanish. Using this value of κ we are then able to fit the parametric plot of E and L to very high accuracy and obtain the classical values to high numerical precision.

We close this section with the remark that the confining behavior appearing in the IR is not something expected from a gauge theory point of view. For this reason, it would be interesting to examine the stability of the string trajectory used for calculating the quark-antiquark potential under small fluctuations, through the tools obtained in [41, 42].

7. Discussion

In this paper we have studied the addition of flavor to the supergravity dual of three-dimensional gauge theories with eight supersymmetries. The unflavored background is constructed by wrapping D4-branes in a two cycle of a Calabi-Yau twofold. The corresponding flavor branes are also D4-branes that extend along the non-compact directions

of the normal bundle of the cycle. We have shown that the addition of these flavor branes does not break any of the eight supersymmetries preserved by the unflavored background. We have studied this system both in the quenched and unquenched approaches and we have obtained the meson spectrum by analyzing the normalizable fluctuations of the probe flavor D4-branes. In the brane probe approximation we have also studied the Higgs branch of the theory, which is realized as a configuration in which the color and flavor D4-branes are recombined.

We have succeeded in computing the backreaction of a large number of flavor branes by considering a continuous distribution of D4-branes smeared over their transverse angular directions. The flavor branes provide a source term for the RR fields that induces a violation of the Bianchi identity for F_4 . The ansatz for the backreacted background can be obtained by modifying the unflavored one to satisfy the new Bianchi identities. By requiring that the flavored background preserves the same amount of supersymmetry as the unflavored solution, one arrives at a system of first-order BPS equations. We have checked that the BPS equations imply the Maxwell-Einstein equations with sources. The functions in the ansatz depend on the two radial variables σ and r and the BPS equations are a system of PDEs that must be integrated numerically. However, the solution for $\sigma = 0$ can be found analytically and we have shown that it reproduces the running of the coupling of the Yang-Mills theory with flavors.

The supergravity solutions we have studied do not capture the rich non-perturbative structure of the corresponding field theory duals. This fact is related to the bad IR behavior of the solutions, which have a naked singularity in this region. This problem is similar to the one presented by other similar backgrounds with the same amount of supersymmetry in four [8, 15] and two dimensions [22]. As in these cases, one can argue that the singularity can be consistently screened by an enhançon, which is the locus where the sources of the background become effectively tensionless and the geometry ends. The fact that the gravity solution reproduces the exact perturbative behavior and misses the non-perturbative effects seems to be due to the suppression of the latter in the 't Hooft large N_c limit.

We have also studied the behavior of Wilson loops in the unflavored background, which critically depends on the value of the parameter κ of the solution. While for $\kappa \leq 1/16$ there is a maximal separation of the quark-antiquark pair, for $\kappa > 1/16$ there is a transition from a Coulomb-like behavior at small separation to a linear potential for large separation. Above a certain critical value of $\kappa = \kappa_{cr}$ the transition is discontinuous (first order) while exactly at the critical value of κ it becomes second order. At this point we have estimated the critical exponents, which are given by the classical mean field theory values. Actually, this critical behavior seems to be universal, at least for these type of models: we have found it in the 4d system of [8], as well as in the 2d system of [22]. In both cases there is a constant which is the analogue of κ and a critical value for this constant. Similar results have been found in [39, 36, 40, 19, 20, 23] in other backgrounds. Interestingly, the critical exponents in these systems are also given by the classical values. To properly interpret these results one should have a clean understanding of the meaning of the constant κ (and of its analogues in other models). The most natural interpretation is that κ parametrizes the VEV of some operator that is switched on. However, more work is needed to confirm

this interpretation and to find the precise nature of the phase transition uncovered. We are working in this direction.

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A. BPS equations

Let us derive the BPS equations (3.12) of the general flavored background by imposing the preservation of eight supersymmetries. For the type of background we are studying, the supersymmetry variations for the dilatino λ and gravitino ψ_M in the type IIA theory in the string frame are:

$$\begin{aligned} \delta\lambda &= \frac{1}{2} \left[\Gamma^M \partial_M \Phi - \frac{1}{4 \cdot 4!} e^\Phi F_{MNPQ}^{(4)} \Gamma^{MNPQ} \right] \epsilon, \\ \delta\psi_M &= \left[\nabla_M - \frac{1}{8 \cdot 4!} e^\Phi F_{NPQR}^{(4)} \Gamma^{NPQR} \Gamma_M \right] \epsilon, \end{aligned} \tag{A.1}$$

where ϵ is a ten-dimensional spinor. The Killing spinors of the background are those for which $\delta\lambda = \delta\psi_M = 0$. They will be characterized by a set of algebraic projection conditions that can be expressed in terms of products of constant Dirac matrices with flat indices. In order to specify these conditions, let us choose the following vielbein basis:

$$\begin{aligned} e^{0,1,2} &= e^\Phi dx^{0,1,2}, & e^3 &= e^\Phi \mathcal{Z}^{\frac{1}{2}} R d\tilde{\theta}, & e^4 &= e^\Phi \mathcal{Z}^{\frac{1}{2}} R \sin\tilde{\theta} d\tilde{\phi}, \\ e^5 &= e^{-\Phi} \mathcal{Z}^{-\frac{1}{2}} d\sigma, & e^6 &= e^{-\Phi} \mathcal{Z}^{-\frac{1}{2}} \sigma \left(d\psi + \cos\tilde{\theta} d\tilde{\phi} \right), \\ e^7 &= e^{-\Phi} dr, & e^8 &= e^{-\Phi} r d\theta, & e^9 &= e^{-\Phi} r \sin\theta d\phi. \end{aligned} \tag{A.2}$$

It is useful to express the four-form (3.11) in terms of the vielbein basis (A.2). One has:

$$\begin{aligned} F_4 &= -\frac{e^{4\Phi}}{\sigma r^2} \sqrt{\mathcal{Z}} \left(g' e^7 + \sqrt{\mathcal{Z}} \dot{g} e^5 \right) \wedge e^6 \wedge e^8 \wedge e^9 + \\ &+ \frac{1}{\mathcal{Z} r^2 R^2} \left[g + \frac{N_f}{2 N_c} \frac{R^3}{8} \Theta(r - r_Q) \right] e^3 \wedge e^4 \wedge e^8 \wedge e^9. \end{aligned} \tag{A.3}$$

Then, the projections satisfied by the Killing spinors are:

$$\Gamma_{01234} \Gamma_{11} \epsilon = \epsilon, \quad \Gamma_{01256} \Gamma_{11} \epsilon = -\epsilon, \tag{A.4}$$

where the indices of the Γ 's refer to the frame basis (A.2). One can easily show that, after putting all these ingredients into the supersymmetry transformations for the gravitino

and dilatino (A.1), we end up with the system (3.12) of first-order BPS equations for the functions of our ansatz. Moreover, the expression for the Killing spinors is:

$$\epsilon = e^{\frac{\Phi}{2}} e^{-\frac{\psi}{2}\Gamma_{34}} e^{\frac{\theta}{2}\Gamma_{78}} e^{\frac{\phi}{2}\Gamma_{89}} \eta, \quad (\text{A.5})$$

where η is a constant spinor that satisfies the same projections (A.4) as ϵ . Notice that these algebraic conditions imply that the system is 1/4-supersymmetric, i.e. that eight supersymmetries are preserved.

Let us now show that the first-order BPS system (3.12) implies the second-order equations of motion of the different fields. First of all, let us start by checking the equation of motion of F_4 , which is given by:

$$d(\star F_4) = 0, \quad (\text{A.6})$$

One can prove that (A.6) is equivalent to the following PDE:

$$\partial_r \left[\frac{\mathcal{Z} g' e^{8\Phi}}{\sigma r^2} \right] + \partial_\sigma \left[\frac{\mathcal{Z}^2 \dot{g} e^{8\Phi}}{\sigma r^2} \right] = \frac{\sigma}{R^4 r^2 \mathcal{Z}^2} \left[g + \frac{N_f}{2 N_c} \frac{R^3}{8} \Theta(r - r_Q) \right]. \quad (\text{A.7})$$

Let us verify that equation (A.7) is satisfied as a consequence of the system (3.12). We will check this fact by an explicit calculation. First, by using (3.12), we rewrite all terms appearing in (A.7) as:

$$\begin{aligned} \frac{\mathcal{Z} g' e^{8\Phi}}{\sigma r^2} &= -\frac{\mathcal{Z}}{R^2} \partial_\sigma \left(\frac{\sigma}{\mathcal{Z} \dot{\mathcal{Z}}} \right), \\ \frac{\mathcal{Z}^2 \dot{g} e^{8\Phi}}{\sigma r^2} &= \frac{1}{R^2} \partial_r \left(\frac{\sigma}{\dot{\mathcal{Z}}} \right), \\ \frac{\sigma}{\mathcal{Z}^2 r^2} \left[g + \frac{N_f}{2 N_c} \frac{R^3}{8} \Theta(r - r_Q) \right] &= R^2 \partial_r \left(\frac{\sigma}{\dot{\mathcal{Z}}} \right). \end{aligned} \quad (\text{A.8})$$

By plugging this result into (A.7) one can straightforwardly verify (A.6). Thus, $F_6 = \star F_4$ should be represented as the derivative of a five-form, dC_5 . Indeed, by using again (3.12) and (A.8) one can readily obtain an expression for this five-form potential, namely:

$$C_5 = dx^0 \wedge dx^1 \wedge dx^2 \wedge \left[R^2 \mathcal{Z} e^{4\Phi} \tilde{\omega}_2 - \frac{\sigma}{\dot{\mathcal{Z}}} d\sigma \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) \right]. \quad (\text{A.9})$$

In order to check that the Einstein and dilaton equations of motion are satisfied in the flavored background, after using the BPS equations, we write them in the Einstein frame. In this frame we will use a one-form basis which is just the one in (A.2) conveniently rescaled with the exponential of the dilaton as:

$$E^{\bar{M}} = e^{-\frac{\Phi}{4}} e^{\bar{M}}. \quad (\text{A.10})$$

Notice that the smearing form Ω of (3.5) can be written in this basis as:

$$\Omega = \frac{N_f}{16\pi^2} \frac{e^{9\Phi/4}}{R^2} \frac{\delta(r - r_Q)}{r^2 \mathcal{Z}} E^3 \wedge E^4 \wedge E^7 \wedge E^8 \wedge E^8. \quad (\text{A.11})$$

Moreover, the DBI action for the smeared flavor branes in the Einstein frame takes the form:

$$S_{\text{DBI}}^E = -T_4 \int_{\mathcal{M}_{10}} d^{10}x e^{\Phi/4} \sqrt{-\det G^E} |\Omega|^E, \quad (\text{A.12})$$

where $G^E = e^{-\Phi/2} G$ is the Einstein frame metric and the modulus $|\Omega|^E$ is computed with G^E (from now on we will suppress the index E of G^E). It follows immediately from (A.11) that $|\Omega|^E$ is given by:

$$|\Omega|^E = \frac{N_f}{16\pi^2} \frac{e^{9\Phi/4}}{R^2} \frac{\delta(r-r_Q)}{r^2 |\mathcal{Z}|}. \quad (\text{A.13})$$

The total action of the system is just the one of type IIA supergravity plus the DBI+WZ action of the flavor branes. The corresponding equation of motion for the dilaton is:

$$\square \Phi = \frac{1}{4 \cdot 4!} e^{\Phi/2} F_4^2 + 2\kappa_{10}^2 T_4 \frac{e^{\Phi/4}}{4} |\Omega|^E. \quad (\text{A.14})$$

Again, one can show that (A.14) is satisfied as a consequence of the system (3.12). The verification of this fact is straightforward (but tedious) and will not be detailed here. Let us just mention the fact that one has to evaluate the left-hand side of (A.14) by computing the derivatives of the first-order BPS equations in (3.12). In this process one generates some terms in which the Heaviside function $\Theta(r-r_Q)$ is differentiated and, therefore, the Dirac delta function $\delta(r-r_Q)$ is produced. These terms match precisely the one containing $|\Omega|^E$ in (A.14), while the remaining ones correspond to the F_4^2 term.

The Einstein equations in the Einstein frame are:

$$R_{MN} - \frac{1}{2} G_{MN} R = \frac{1}{2} \left[\partial_M \Phi \partial_N \Phi - \frac{1}{2} G_{MN} (\partial \Phi)^2 \right] + \frac{1}{2 \cdot 4!} e^{\Phi/2} \left[4(F_4^2)_{MN} - \frac{1}{2} G_{MN} F_4^2 \right] + T_{MN}, \quad (\text{A.15})$$

where T_{MN} is the energy-momentum tensor of the smeared flavor brane, defined as:

$$T_{MN} = -\frac{2\kappa_{10}^2}{\sqrt{-G}} \frac{\delta S_{\text{DBI}}^E}{\delta G^{MN}}. \quad (\text{A.16})$$

Taking into account the form of S_{DBI} in (A.12) we arrive at the following expression for T_{MN} in flat components:

$$T_{\bar{M}\bar{N}} = 2\kappa_{10}^2 T_4 \frac{e^{\Phi/4}}{2} \left[\eta_{\bar{M}\bar{N}} |\Omega|^E - \frac{1}{4!} \frac{1}{|\Omega|^E} (\Omega^2)_{\bar{M}\bar{N}} \right]. \quad (\text{A.17})$$

From this expression one readily gets the explicit values of the different components of $T_{\bar{M}\bar{N}}$, which are:

$$\begin{aligned} -T_{00} = T_{11} = T_{22} = T_{88} = T_{99} &= \frac{R}{8} \frac{N_f}{2N_c} e^{5\Phi/2} \frac{\delta(r-r_Q)}{2r^2 \mathcal{Z}}, \\ T_{33} = T_{44} = T_{55} = T_{66} = T_{77} &= 0. \end{aligned} \quad (\text{A.18})$$

By using these values one can verify, following the same strategy used to prove (A.14), that the Einstein equations (A.15) are satisfied as a consequence of the first-order equations (3.12).

B. Supersymmetric embeddings

In this appendix we will characterize a class of supersymmetric embeddings of D4-branes in the backgrounds of sections 2 and 3. Our main tool will be kappa symmetry [43], which states that the supersymmetric embeddings of the D4-brane are those that satisfy the condition $\Gamma_\kappa \epsilon = \epsilon$, where ϵ is a Killing spinor of the background and Γ_κ is a matrix which depends on the embedding. To write the precise form of Γ_κ , let us define the induced Dirac matrices on the D4-brane worldvolume as $\gamma_a = \partial_a X^M E_M^{\tilde{M}} \Gamma_{\tilde{M}}$, where $X^M(\xi^a)$ are the functions that parameterize the embedding and $E_M^{\tilde{M}}$ are the vielbein coefficients of ten-dimensional metric. Then, when the worldvolume gauge field F is zero, the matrix Γ_κ for the D4-brane is [44]:

$$\Gamma_\kappa = \frac{1}{5!} \frac{1}{\sqrt{-\det \hat{G}_5}} \Gamma_{11} \epsilon^{a_1 \dots a_5} \gamma_{a_1 \dots a_5} , \quad (\text{B.1})$$

where $\gamma_{a_1 \dots a_5}$ denotes the antisymmetrized product of the induced matrices, \hat{G}_5 is the induced metric on the D4-brane worldvolume and Γ_{11} is the chiral matrix in ten dimensions. Let us choose, as in (3.1), x^0, x^1, x^2, σ and ψ as worldvolume coordinates and let us consider embeddings as in (4.11), in which $\tilde{\theta}$ and $\tilde{\phi}$ depend on (σ, ψ) and the remaining coordinates are constant. Then, the kappa symmetry matrix (B.1) takes the following form:

$$\Gamma_\kappa = \frac{1}{\sqrt{-\det \hat{G}_5}} \Gamma_{11} \gamma_{x^0 x^1 x^2 \sigma \psi} , \quad (\text{B.2})$$

where the induced Gamma matrices are:

$$\begin{aligned} \gamma_{x^0,1,2} &= e^\Phi \Gamma_{0,1,2} , \\ \gamma_\sigma &= R e^\Phi \mathcal{Z}^{1/2} \left[\partial_\sigma \tilde{\theta} \Gamma_3 + \sin \tilde{\theta} \partial_\sigma \tilde{\phi} \Gamma_4 \right] + \frac{1}{e^\Phi \mathcal{Z}^{1/2}} \left[\Gamma_5 + \sigma \cos \tilde{\theta} \partial_\sigma \tilde{\phi} \Gamma_6 \right] , \\ \gamma_\psi &= R e^\Phi \mathcal{Z}^{1/2} \left[\partial_\psi \tilde{\theta} \Gamma_3 + \sin \tilde{\theta} \partial_\psi \tilde{\phi} \Gamma_4 \right] + \frac{\sigma}{e^\Phi \mathcal{Z}^{1/2}} \left[1 + \cos \tilde{\theta} \partial_\psi \tilde{\phi} \right] \Gamma_6 . \end{aligned} \quad (\text{B.3})$$

To find those embeddings that are kappa symmetric and preserve the same amount of supersymmetry as the original background, we should compute the action of the antisymmetrized product $\gamma_{x^0 x^1 x^2 \sigma \psi}$ on the spinor. Taking into account the projections (A.4), we have:

$$e^{-3\Phi} \Gamma_{11} \gamma_{x^0 x^1 x^2 \sigma \psi} \epsilon = [c_I + c_{39} \Gamma_{39} + c_{49} \Gamma_{49}] \epsilon , \quad (\text{B.4})$$

where the coefficients appearing on the right-hand side of the above equation are:

$$\begin{aligned} c_I &= \frac{\sigma}{e^{2\Phi} \mathcal{Z}} \left[1 + \sigma \cos \tilde{\theta} \partial_\psi \tilde{\phi} \right] + R^2 e^{2\Phi} \mathcal{Z} \sin \tilde{\theta} \left[\partial_\psi \tilde{\theta} \partial_\sigma \tilde{\phi} - \partial_\sigma \tilde{\theta} \partial_\psi \tilde{\phi} \right] , \\ c_{39} &= R \partial_\psi \tilde{\theta} - R \sigma \sin \tilde{\theta} \partial_\sigma \tilde{\phi} , \\ c_{49} &= R \sin \tilde{\theta} \partial_\psi \tilde{\phi} + R \sigma \left[\partial_\sigma \tilde{\theta} + \cos \tilde{\theta} \left(\partial_\psi \tilde{\phi} \partial_\sigma \tilde{\theta} - \partial_\sigma \tilde{\phi} \partial_\psi \tilde{\theta} \right) \right] . \end{aligned} \quad (\text{B.5})$$

The right-hand side of (B.4) should contain only the term with the identity matrix if we want to satisfy the $\Gamma_\kappa \epsilon = \epsilon$ condition for any Killing spinor of the background. Thus, we must demand that:

$$c_{39} = c_{49} = 0 , \quad (\text{B.6})$$

which leads us to the following system of PDE's:

$$\begin{aligned} \partial_\psi \tilde{\theta} - \sigma \sin \tilde{\theta} \partial_\sigma \tilde{\phi} &= 0, \\ \sin \tilde{\theta} \partial_\psi \tilde{\phi} + \sigma \left[\partial_\sigma \tilde{\theta} + \cos \tilde{\theta} \left(\partial_\psi \tilde{\phi} \partial_\sigma \tilde{\theta} - \partial_\sigma \tilde{\phi} \partial_\psi \tilde{\theta} \right) \right] &= 0. \end{aligned} \tag{B.7}$$

The general solution of the system (B.7) was obtained in ref. [22]. As mentioned in subsection 4.2, if one defines the two complex variables ζ_1 and ζ_2 as in (4.12), any holomorphic function $\zeta_1 = f(\zeta_2)$ solves (B.7). The corresponding embedding preserves the eight supersymmetries of the background.

C. Additional supergravity backgrounds

In this appendix we find additional supergravity solutions for our unflavored setup. First of all we will show that it is possible to find solutions for the unflavored BPS system of equations (2.8) that are simpler than those coming from gauged supergravity. Later on, in subsection C.1, after performing a series of duality transformations to the unflavored solution of section 2, we will generate a background dual to a non-relativistic system in 1+1 dimensions.

Let us try to solve the unflavored BPS system (2.8) by means of the method of separation of variables. Accordingly, let us adopt the following ansatz for the three functions g , Φ and \mathcal{Z} appearing in the background:

$$g = g(\sigma), \quad \Phi = \Phi(r), \quad \mathcal{Z} = \mathcal{Z}_1(r) \mathcal{Z}_2(\sigma). \tag{C.1}$$

It is straightforward to show that the general solution of the system (2.8) that has the form of the ansatz (C.1) is:

$$\begin{aligned} g(\sigma) &= c_1 R^2 \sqrt{c_4 + c_3 \sigma^2}, \\ e^{-4\Phi(r)} &= c_3 R^2 \left[c_2 + \frac{c_1}{r} \right]^2, \\ \mathcal{Z}(r, \sigma) &= \left(c_2 + \frac{c_1}{r} \right) \sqrt{c_4 + c_3 \sigma^2}, \end{aligned} \tag{C.2}$$

where the c_i 's are separation constants. The ten-dimensional metric corresponding to this solution can be obtained by plugging the values of $e^{-4\Phi}$ and \mathcal{Z} given in (C.2) into a general ansatz. Proceeding in this way we arrive at the split 6+4 metric:

$$ds_{10}^2 = ds_6^2 + ds_4^2, \tag{C.3}$$

where the six-dimensional metric is independent of σ :

$$ds_6^2 = e^{2\Phi} dx_{1,2}^2 + e^{-2\Phi} \left[dr^2 + r^2 d\Omega_2^2 \right], \tag{C.4}$$

while the four-dimensional metric is independent of r :

$$ds_4^2 = \frac{R \sqrt{c_4 + c_3 \sigma^2}}{\sqrt{c_3}} \left(d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2 \right) + \frac{R \sqrt{c_3}}{\sqrt{c_4 + c_3 \sigma^2}} \left[d\sigma^2 + \sigma^2 \left(d\psi + \cos \tilde{\theta} d\tilde{\phi} \right)^2 \right]. \tag{C.5}$$

In order to rewrite the metric (C.5) in a more familiar form, let us perform a change of variables and define a new radial variable ζ , related to σ in the following way:

$$\zeta^2 = \frac{4R}{\sqrt{c_3}} \sqrt{c_4 + c_3 \sigma^2} . \quad (\text{C.6})$$

Actually, if we define a new constant a as:

$$a^4 = \frac{16R^2 c_4}{c_3} , \quad (\text{C.7})$$

the relation that gives σ in terms of ζ is:

$$\sigma = \frac{1}{4R} \sqrt{\zeta^4 - a^4} . \quad (\text{C.8})$$

Clearly, $\zeta \geq a$, which corresponds to the range $\sigma \geq 0$. After this change of variable, one can easily prove that the metric (C.5) becomes:

$$ds_4^2 = \frac{d\zeta^2}{1 - \left(\frac{a}{\zeta}\right)^4} + \frac{\zeta^2}{4} \left[d\tilde{\Omega}_2^2 + \left(1 - \left(\frac{a}{\zeta}\right)^4\right) (d\psi + \cos \tilde{\theta} d\tilde{\phi})^2 \right] , \quad (\text{C.9})$$

which is the metric of an Eguchi-Hanson space EH_4 with resolution parameter a . When $a = 0$ the metric (C.9) becomes the one corresponding to the $\mathbb{C}^2/\mathbb{Z}_2$ orbifold. Moreover, if we define the new constants η and Q as:

$$\eta \equiv R \sqrt{c_3} c_2 , \quad Q \equiv R \sqrt{c_3} c_1 , \quad (\text{C.10})$$

then, the warp factor $e^{-4\Phi}$ becomes:

$$e^{-4\Phi} = \left(\eta + \frac{Q}{r} \right)^2 . \quad (\text{C.11})$$

Using this result the six-dimensional part of the metric takes the form:

$$ds_6^2 = \frac{dx_{1,2}^2}{\eta + \frac{Q}{r}} + \left(\eta + \frac{Q}{r} \right) \left[dr^2 + r^2 d\Omega_2^2 \right] . \quad (\text{C.12})$$

Notice that, for $\eta \neq 0$, the metric (C.12) is asymptotically flat.

C.1 Non-relativistic backgrounds

Let us now follow the procedure of ref. [45] to obtain backgrounds dual to non-relativistic systems by performing a combination of two T-dualities and a shift to a supergravity dual of a relativistic theory. First of all, we introduce light-cone variables in the standard way, $x^\pm = x^0 \pm x^1$, and rewrite the initial metric and the dilaton of the unflavored type IIA background as:

$$\begin{aligned} ds_{\text{IIA}}^2 &= H^{-1/2} \left[-dx^+ dx^- + (dx^2)^2 + \mathcal{Z} R^2 \left(d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2 \right) \right] + \\ &\quad + H^{1/2} \left[dr^2 + r^2 d\Omega_2^2 + \frac{1}{\mathcal{Z}} \left(d\sigma^2 + \sigma^2 \left(d\psi + \cos \tilde{\theta} d\tilde{\phi} \right)^2 \right) \right] , \\ e^{2\Phi} &= H^{-1/2} . \end{aligned} \quad (\text{C.13})$$

Notice that we have introduced the warp factor H . Similarly, the RR potentials in these variables take the form:

$$\begin{aligned} C_3 &= -g \omega_2 \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) , \\ C_5 &= -\frac{1}{2} dx^+ \wedge dx^- \wedge dx^2 \wedge \left[\frac{R^2 \mathcal{Z}}{H} \tilde{\omega}_2 - \frac{\sigma}{\mathcal{Z}} d\sigma \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) \right] . \end{aligned} \quad (\text{C.14})$$

In the first step we perform a T-duality along the fiber direction ψ . In general, under T-duality the fields in the NSNS sector form a closed set and transform amongst themselves. Hence for these fields we may use the standard rules. In contrast, the transformation rules for the RR sector fields involve those in the NSNS sector. The metric and dilaton after this T-duality become:

$$\begin{aligned} ds_{\text{IIB}}^2 &= H^{-1/2} \left[-dx^+ dx^- + (dx^2)^2 + \mathcal{Z} R^2 \left(d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2 \right) \right] + \\ &\quad + H^{1/2} \left[dr^2 + r^2 d\Omega_2^2 + \frac{1}{\mathcal{Z}} d\sigma^2 \right] + H^{-1/2} \frac{\mathcal{Z}}{\sigma^2} d\psi^2 , \\ e^{2\Phi} &= \frac{\mathcal{Z}}{\sigma^2} H^{-1} , \end{aligned} \quad (\text{C.15})$$

while the non-vanishing NSNS and RR potentials are:

$$\begin{aligned} B_2 &= \cos \tilde{\theta} d\psi \wedge d\tilde{\phi} , \\ C_2 &= -g\omega_2 , \quad C_4 = \frac{\sigma}{2\mathcal{Z}} dx^+ \wedge dx^- \wedge dx^2 \wedge d\sigma , \\ C_6 &= -\frac{R^2 \mathcal{Z}}{2H} dx^+ \wedge dx^- \wedge dx^2 \wedge \tilde{\omega}_2 \wedge d\psi . \end{aligned} \quad (\text{C.16})$$

In the second step we perform a coordinate shift along the light-cone coordinate x^- of the form:

$$x^- \rightarrow x^- + \gamma\psi \quad \longrightarrow \quad dx^- \rightarrow dx^- + \gamma d\psi . \quad (\text{C.17})$$

The corresponding metric and the dilaton become:

$$\begin{aligned} ds_{\text{IIB}}^2 &= H^{-1/2} \left[-dx^+ dx^- - \gamma dx^+ d\psi + (dx^2)^2 + \mathcal{Z} R^2 \left(d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2 \right) \right] + \\ &\quad + H^{1/2} \left[dr^2 + r^2 d\Omega_2^2 + \frac{1}{\mathcal{Z}} d\sigma^2 \right] + H^{-1/2} \frac{\mathcal{Z}}{\sigma^2} d\psi^2 , \\ e^{2\Phi} &= \frac{\mathcal{Z}}{\sigma^2} H^{-1} . \end{aligned} \quad (\text{C.18})$$

while the non-vanishing NSNS and RR potentials are:

$$\begin{aligned} B_2 &= \cos \tilde{\theta} d\psi \wedge d\tilde{\phi} , \\ C_2 &= -g\omega_2 , \quad C_4 = \frac{\sigma}{2\mathcal{Z}} dx^+ \wedge (dx^- + \gamma d\psi) \wedge dx^2 \wedge d\sigma , \\ C_6 &= -\frac{R^2 \mathcal{Z}}{2H} dx^+ \wedge dx^- \wedge dx^2 \wedge \tilde{\omega}_2 \wedge d\psi . \end{aligned} \quad (\text{C.19})$$

Finally we perform another T-duality along ψ and return to a type IIA background. Now the metric and the dilaton become:

$$\begin{aligned}
 ds_{\text{IIA}}^2 &= H^{-1/2} \left[-dx^+ \left(dx^- + \frac{\gamma^2 \sigma^2}{4 \mathcal{Z}} dx^+ \right) + (dx^2)^2 + \mathcal{Z} R^2 \left(d\tilde{\theta}^2 + \sin^2 \tilde{\theta} d\tilde{\phi}^2 \right) \right] + \\
 &\quad + H^{1/2} \left[dr^2 + r^2 d\Omega_2^2 + \frac{1}{\mathcal{Z}} \left(d\sigma^2 + \sigma^2 \left(d\psi + \cos \tilde{\theta} d\tilde{\phi} \right)^2 \right) \right], \\
 e^{2\Phi} &= H^{-1/2}.
 \end{aligned} \tag{C.20}$$

while the non-vanishing NSNS and RR potentials are:

$$\begin{aligned}
 B_2 &= \frac{\gamma \sigma^2}{2 \mathcal{Z}} dx^+ \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}), \\
 C_3 &= -g \omega_2 \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) + \frac{\gamma \sigma}{2 \mathcal{Z}} dx^+ \wedge dx^2 \wedge d\sigma, \\
 C_5 &= -\frac{1}{2} dx^+ \wedge dx^- \wedge dx^2 \wedge \left[\frac{R^2 \mathcal{Z}}{2H} \tilde{\omega}_2 - \frac{\sigma}{\mathcal{Z}} d\sigma \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) \right].
 \end{aligned} \tag{C.21}$$

The corresponding field strengths of these potentials are $H_3 = dB_2$, $F_4 = dC_3$. By computing the exterior derivatives and using the BPS equations we have:

$$\begin{aligned}
 H_3 &= \frac{\gamma}{2} \left[\frac{g\sigma^2}{\mathcal{Z}^2 r^2 R^2} dr + \frac{2\sigma}{\mathcal{Z}} \left(1 - \frac{\sigma^2 H}{2\mathcal{Z}^2 R^2} d\sigma \right) \right] \wedge dx^+ \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) \\
 &\quad + \frac{\gamma \sigma^2}{2 \mathcal{Z}} dx^+ \wedge \tilde{\omega}_2, \\
 F_4 &= -dg \wedge \omega_2 \wedge (d\psi + \cos \tilde{\theta} d\tilde{\phi}) + g \omega_2 \wedge \tilde{\omega}_2 + \frac{\gamma}{2} \frac{\sigma g}{r^2 \mathcal{Z}^2 R^2} dx^+ \wedge dx^2 \wedge dr \wedge d\sigma.
 \end{aligned} \tag{C.22}$$

As a consistency check we can verify that these forms satisfy their equations of motion:

$$d\left(e^{-2\Phi} * H_3\right) = \frac{1}{2} F_4 \wedge F_4, \quad d\left(* F_4\right) = H_3 \wedge F_4. \tag{C.23}$$

D. Entanglement entropy

In quantum field theory the entanglement entropy between two complementary spatial regions A and B is defined as the entropy seen by an observer in A who does not have access to the degrees of freedom of B . The holographic computation consists of finding the eight-dimensional surface Σ with minimal area such that its boundary coincides with the boundary of A . Then, the entanglement entropy between A and its complementary region B is given by the integral [46]:

$$S = \frac{1}{4G_{10}} \int_{\Sigma} d^8 \xi e^{-2\Phi} \sqrt{\hat{G}_8}, \tag{D.1}$$

where G_{10} is the ten-dimensional Newton constant, given by $G_{10} = 8\pi^6 \alpha'^4 g_s^2$ and \hat{G}_8 is the induced metric on Σ (see also [47]). We will consider a constant time surface Σ , obtained by minimizing S over all surfaces that approach the boundary of A at the boundary of the ten-dimensional bulk manifold and that are extended along the remaining spatial directions.

The surface we consider is not going to touch the IR region of the space, so it is feasible to use the UV expression (2.24). We parametrize the eight-dimensional surface in the following way:

$$\xi^a = (x, x_2, \tilde{\theta}, \tilde{\phi}, \hat{\alpha}, \theta, \phi, \psi), \quad (\text{D.2})$$

with $u = u(x)$. By computing the induced metric we end up with the following expression for S :

$$S = \frac{2\pi^3 R^5}{3G_{10}} \int_{-\frac{l}{2}}^{\frac{l}{2}} dx u \sqrt{u'^2 + \frac{8}{\sqrt{Z_*}} \left(\frac{u}{R}\right)^3}. \quad (\text{D.3})$$

Since the function S does not depend explicitly on x , the Euler-Lagrange equation derived from S can be integrated with the result being the following:

$$\frac{u^4}{\sqrt{u'^2 + \frac{8}{\sqrt{Z_*}} \left(\frac{u}{R}\right)^3}} = u_0^{5/2} \frac{R^{3/2} Z_*^{1/4}}{2\sqrt{2}}, \quad (\text{D.4})$$

where u_0 is the minimal value of u . From this expression we can obtain u' as a function of u :

$$u' = \pm \sqrt{\frac{8}{R^3 Z_*^{1/2}}} u^{3/2} \sqrt{\frac{u^5}{u_0^5} - 1}. \quad (\text{D.5})$$

Now we can compute the length l as a function of the turning point u_0 of the holographic coordinate:

$$l = 2 \int_{u_0}^{\infty} \frac{du}{|u'|} = \sqrt{2\pi\sqrt{Z_*} R^3} \frac{\Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})} \frac{1}{\sqrt{u_0}}. \quad (\text{D.6})$$

We can use (D.5) to eliminate u' in the entropy functional S . The resulting integral is divergent if the upper limit is infinity. For this reason we regulate it by integrating up to some value u_∞ of u and we have:

$$S = \frac{2\pi^3 R^5}{3G_{10}} u_0^2 \int_1^{\frac{u_\infty}{u_0}} \frac{\xi^{7/2}}{\sqrt{\xi^5 - 1}} d\xi. \quad (\text{D.7})$$

The integral appearing on the right hand side of the above equation takes the value:

$$\int_1^{\frac{u_\infty}{u_0}} \frac{\xi^{7/2}}{\sqrt{\xi^5 - 1}} d\xi = -\frac{\sqrt{\pi}}{2} \frac{\Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})} - \frac{1}{5} B \left[\frac{u_0^5}{u_\infty^5}, -\frac{2}{5}, \frac{1}{2} \right], \quad (\text{D.8})$$

where B is the symbol for the incomplete Beta function. Taking the limit of the above expression when $u_\infty \rightarrow \infty$ we obtain finite and divergent terms:

$$\int_1^{\frac{u_\infty}{u_0}} \frac{\xi^{7/2}}{\sqrt{\xi^5 - 1}} d\xi \approx -\frac{\sqrt{\pi}}{2} \frac{\Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})} + \frac{1}{2} \left(\frac{u_\infty}{u_0}\right)^2 + \mathcal{O}\left(\frac{u_0}{u_\infty}\right)^2. \quad (\text{D.9})$$

Plugging the finite part of this result in (D.7) and using (D.6) to express u_0 in terms of l we have:

$$S^{\text{finite}}(l) = -(2\pi R_{10}) V_\infty \frac{2^5 \sqrt{\pi}}{3} \left[\frac{\Gamma(\frac{3}{5})}{\Gamma(\frac{1}{10})} \right]^5 \frac{N_c^3}{l^4}, \quad (\text{D.10})$$

where $R_{10} = g_s \alpha'^{1/2}$ and V_∞ is the volume of a sphere of radius $RZ_*^{1/2}$ along which the D4-branes are wrapped:

$$V_\infty = 4\pi R^2 Z_* . \quad (\text{D.11})$$

The expression (D.10) just obtained coincides with the one found in [47] for an M5-brane compactified in a two sphere of volume V_∞ . This result was, of course, to be expected, given the UV behavior of our metric.

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